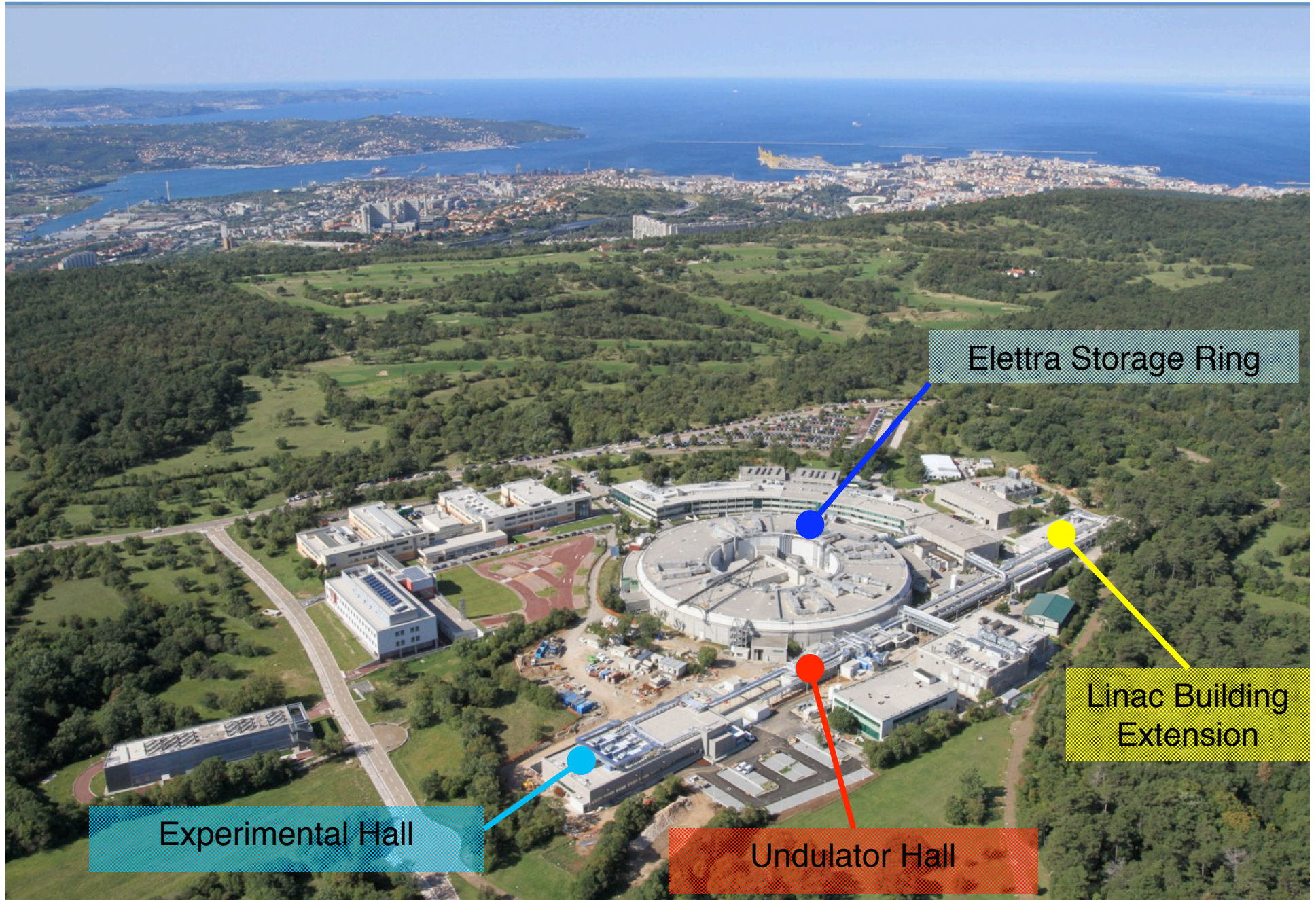


# **First results on the photon beam properties of the seeded FERMI@Elettra FEL1**

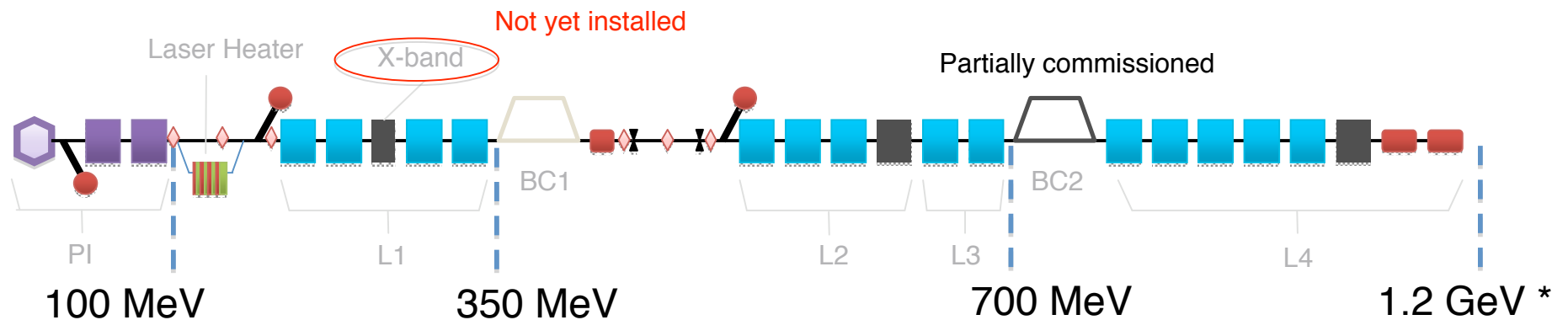
Fulvio Parmigiani  
on behalf of the FERMI@Elettra project

# Overview of Elettra and FERMI





# LINAC Layout

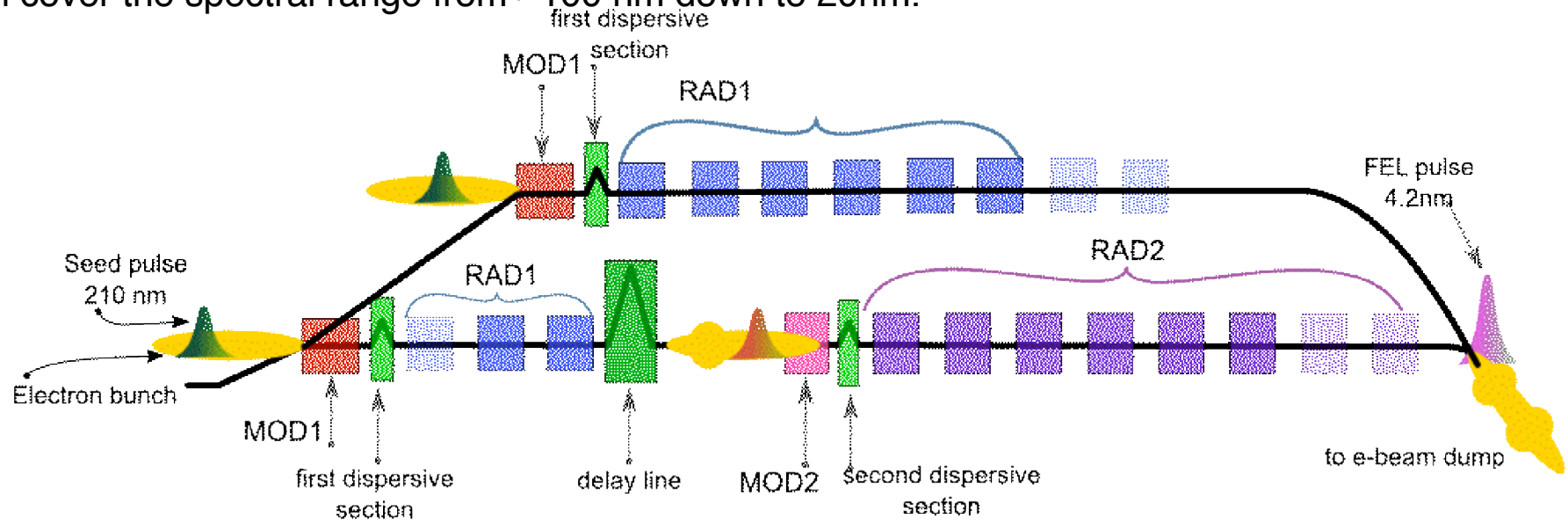


- ❑ Optimized a **450pC - 5ps flat top bunch** at the photo-injector ( $\gamma\epsilon_n=0.9 \mu\text{m}$ )
- ❑ Some test on the Laser Heater but it will be optimized for FEL 2 operation
- ❑ X-band cavity has not yet been installed
- ❑ Beam compressed @BC1 by about **a factor 5** (BC2 not used up to now)
- ❑ Nominal energy @ linac end: **1.2GeV** (reached 1.35GeV)

# FEL-1 and FEL-2

Two FELs will cover different spectral regions.

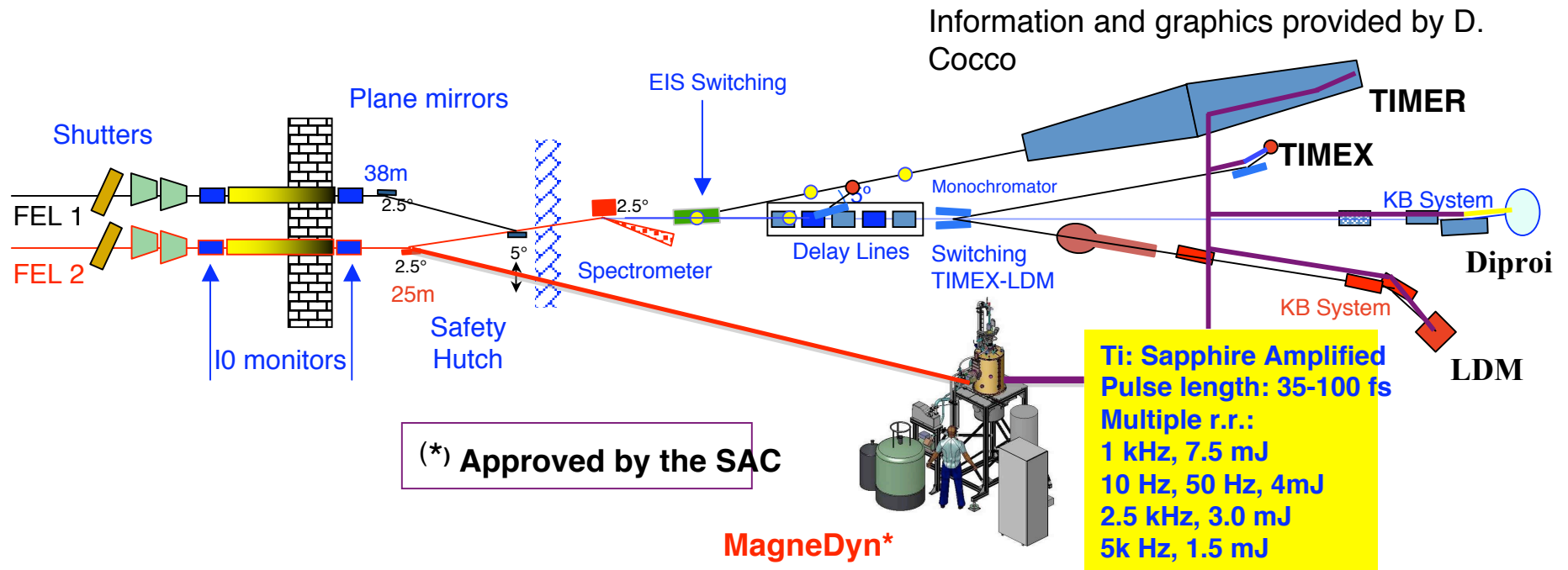
FEL-1, based on a single stage high gain harmonic generations scheme initialized by a UV laser will cover the spectral range from  $\sim 100$  nm down to 20nm.



FEL-2, in order to be able to reach the wavelength range from 20 to  $\sim 4$  nm starting from a seed laser in the UV, will be based on a double cascade of high gain harmonic generation. The nominal layout will use a magnetic electron delay line in order to improve the FEL performance by using the fresh bunch technique. Other FEL configurations are also possible.



# FERMI beamlines-end-stations and pump-laser



LASER SYSTEM: ordered  
LASER HUTCH: under tender  
LASER beam transport to the end-stations in progress

Human resources for the experimental hall (beside beam-line scientists) :

- Experimental hall coordinator
- LASER responsible for the pump-probe experiments

## SCIENCE CASE

---

### } **Low Density Matter** (*coord. C.*

*Callegari*):

- .....brightness
- } structure of nano-clusters .....narrow bw, circular polarization
- } ionization dynamics .....circular polarization
- } magnetism in nano-particles ..... $\lambda$ -tunability
- } catalysis in nano-materials

### } **Elastic and Inelastic Scattering** (*coord. C.*

*Masciovecchio*):

- } Transient Grating Spectroscopy (collective low Fourier Transform Limit dynamics at the nano-scale)
- } Pump & Probe Spectroscopy (meta-stable states of matter) .....brightness,  $\lambda$ -tunability

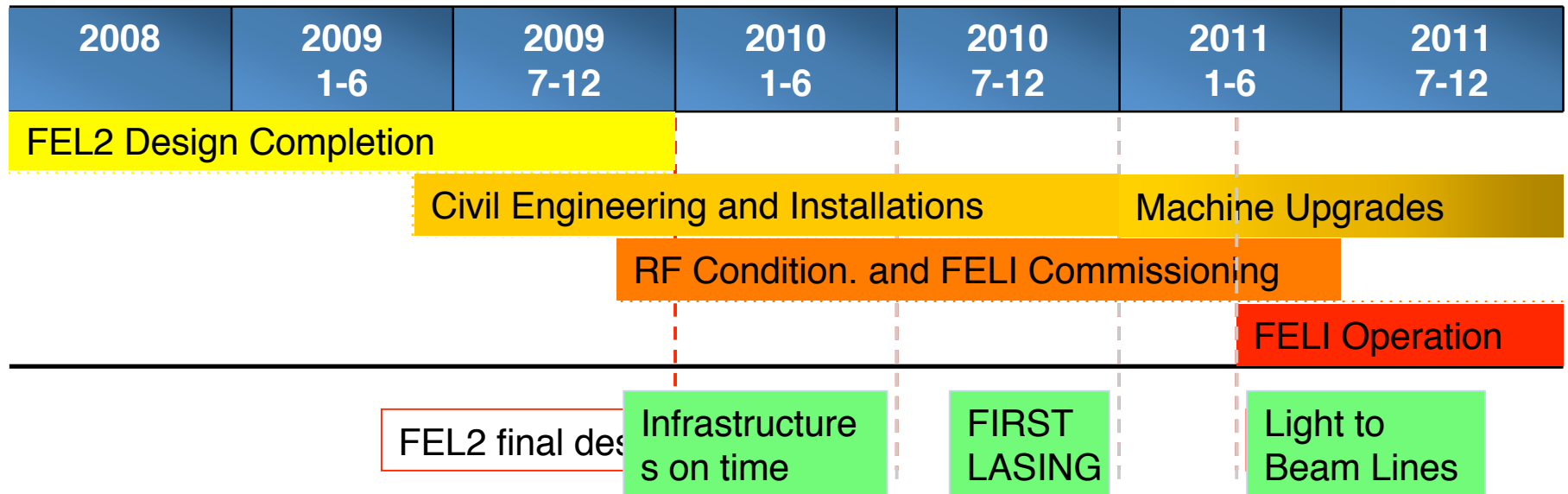
### } **Diffraction and Projection Imaging** (*coord. M. Kiskinova*):

Single-shot & Resonant Transverse Coherent Diffraction Imaging

- } morfology and internal structure at the nm scale .....brightness
- } chemical and magnetic imaging



## DESIGN GOALS & ACHIEVEMENTS



	Parameter	FEL1	FEL2	Units
$\gamma$	Output Wavelength (fund.)	100 – <b>20</b>	20 – 4	nm
	Peak Power	<b>1</b> – 5	> 0.3	GW
	Repetition Rate	<b>10</b>	50	Hz
e	Energy	<b>1.2</b>	1.5	GeV
	Peak Current (core)	<b>200</b> – 800	800	A
	Bunch Length (fhw)	<b>0.7 – 1.2</b>	0.7	ps
	Slice Norm. Emittance	1.5 – <b>3.0</b>	1.0	mm mrad
	Slice Energy Spread	0.20	0.15	MeV

**\* achieved**

# LINAC

The existing 9-structures S-band Linac has been upgraded with:

1. RF photo-cathode Gun (SLAC/BNL/UCLA)  
MeV

$\epsilon_n = 1 \mu\text{m}$  measured at 400pC, 5 ps, 100

2. 7 more CERN/LIL structures

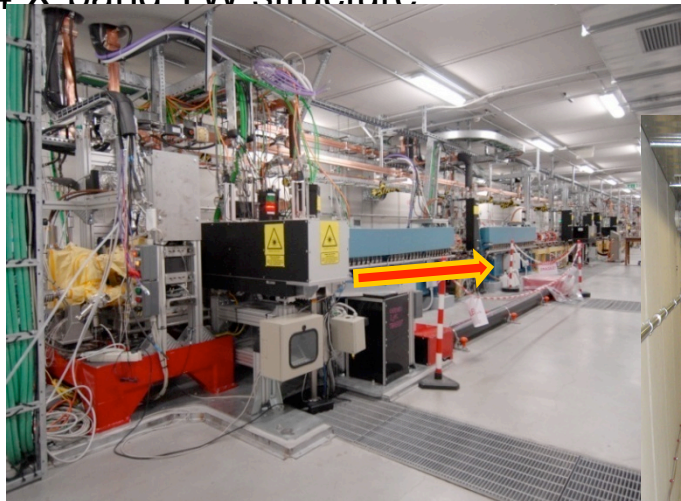
1.35 GeV routinely achieved

3. SLED optimization and phase-modulation

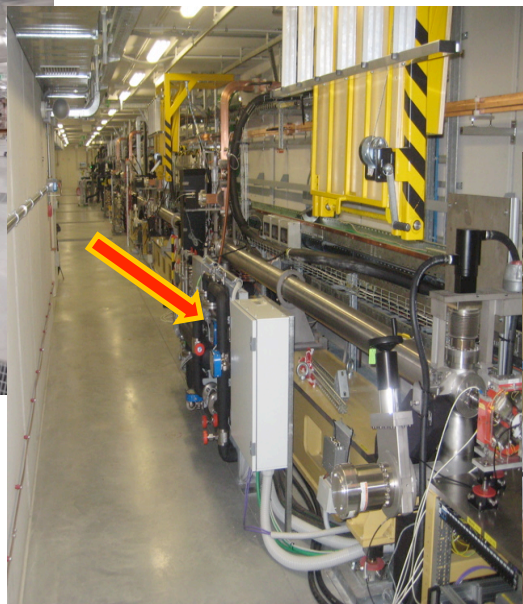
27 MV/m reached for 1.5 GeV operation

4. X-band TW structure

*to be installed for linear time-compression*



PC Gun + Injector



CERN structures



"SLED" structures



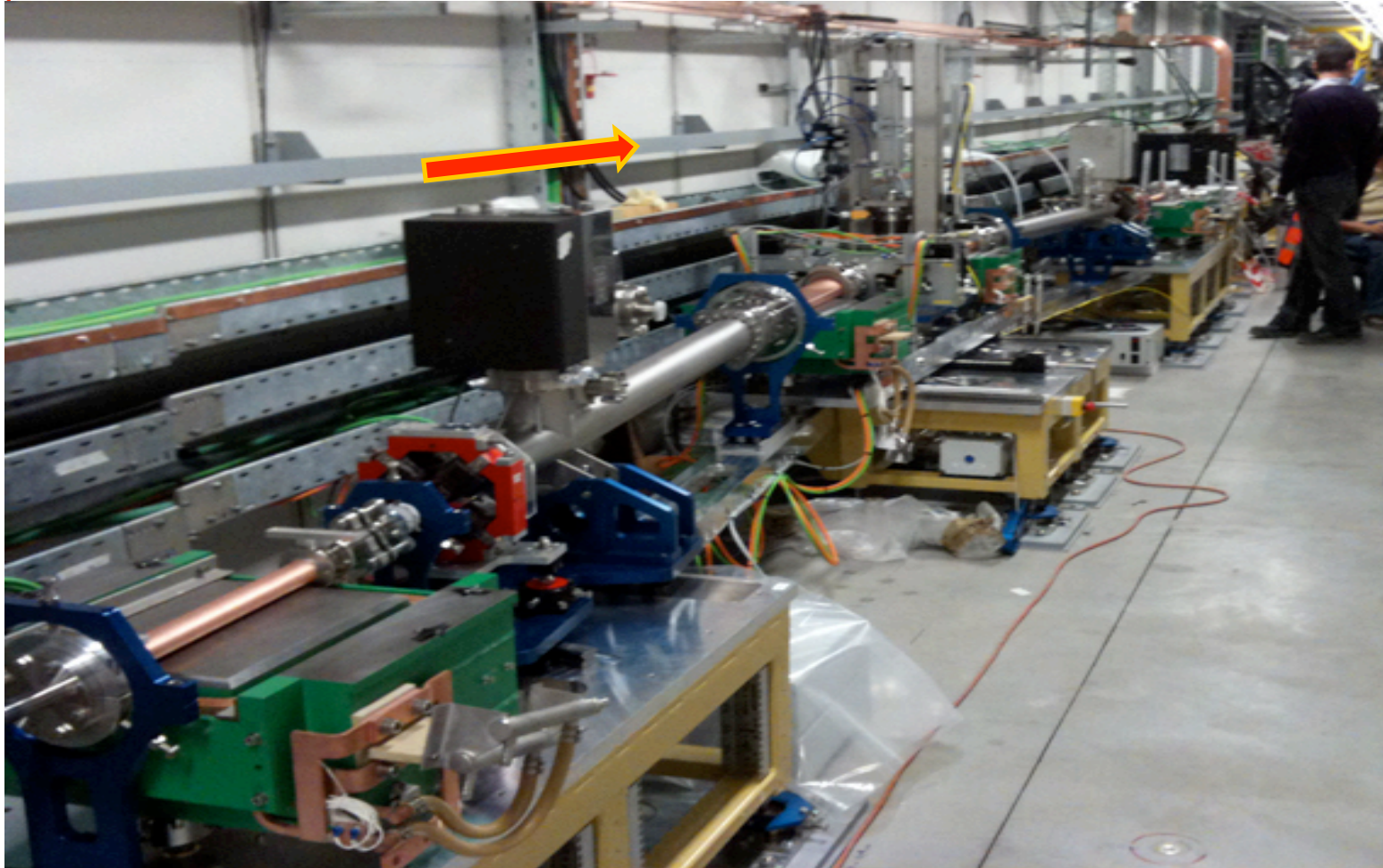
# MAGNETIC COMPRESSOR

---

2 movable magnetic chicanes for one- or two-stage bunch length compression:

1. Developed *in house* on improved LCLS design  
operation

$CF \leq 6$  used for FEL



D. Zangrando (Area Leader), D. La Civita, D. Castronovo, G. Pagon

# TRANSFER LINE

~30 m long high energy transfer line, switching FELI/FEL2. e-Beam diagnostics and collimation included.

It is followed by the undulators (~20 m) and the main dump line (~40 m).



“Spreader”



Main Beam Dump

E. Karantzoulis (Area Leader), S. Ferry, I. Cudin, M. Tudor, et al.

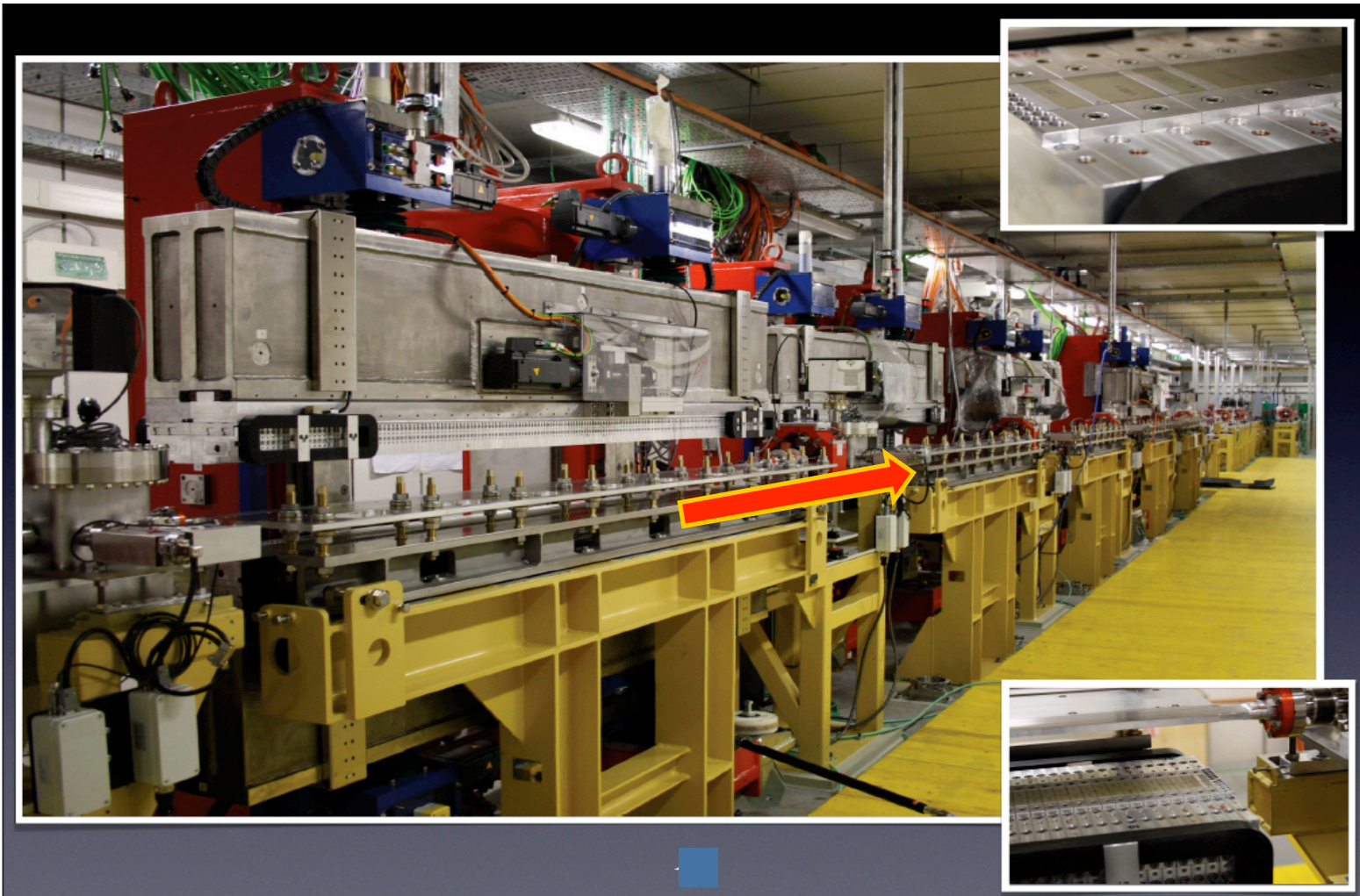


# UNDULATOR S

Variable gap, planar and APPLE-II type Insertion Devices design and manufacturing:

1. Developed *in house* (KYMA spin-off)  
users

variable polarization,  $\lambda$ -tuning provided to



B. Diviacco (Area Leader), D. La Civita, M. Musardo



# PHOTON TRANSPORT SYSTEM

---



D. Cocco (Area Leader), M. Zangrando, C. Svetina

# FEL

S

SASE  
SEEDED  
HGHG  
HHG  
NEW SCHEME

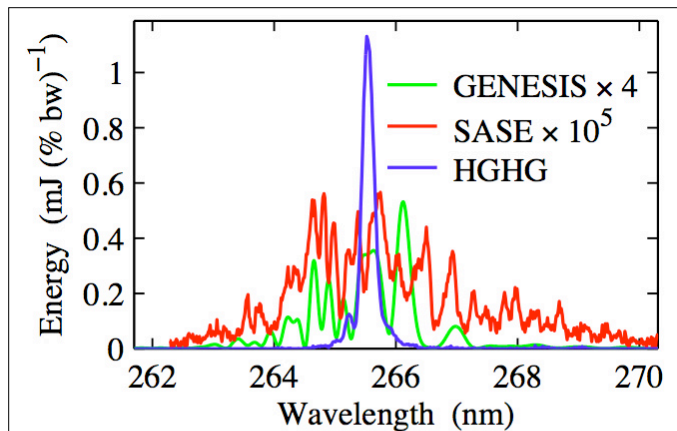
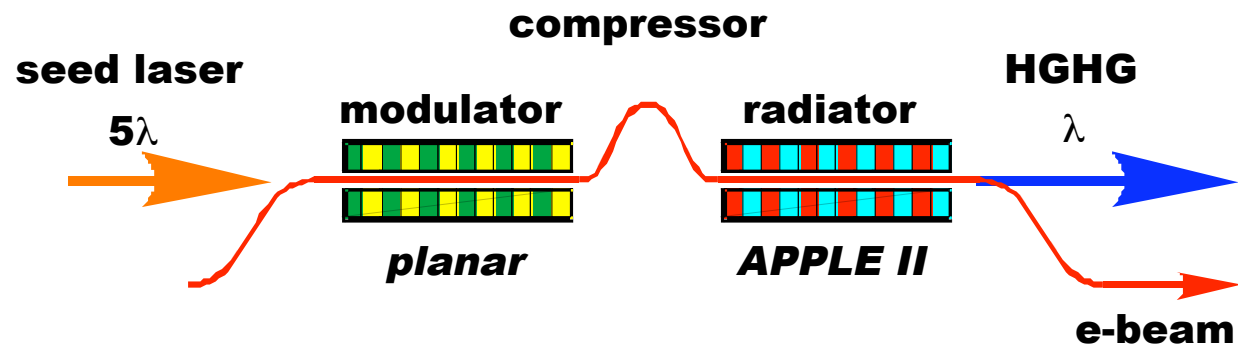
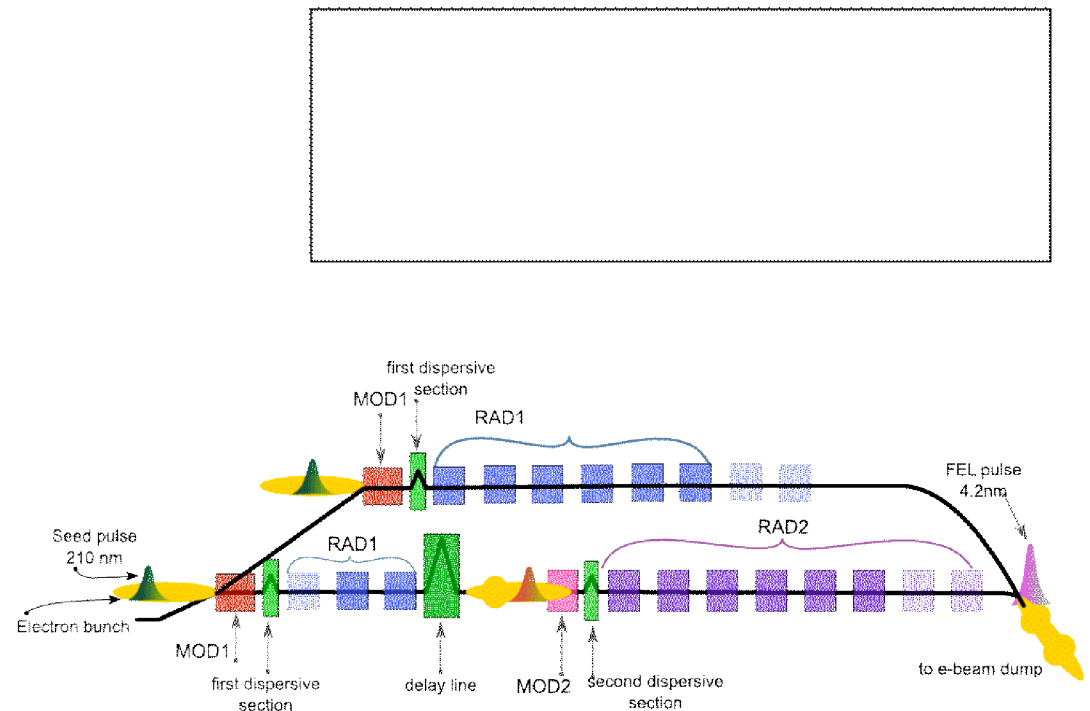


FIG. 4: Single shot HGHG spectrum for 30 MW seed (blue), single shot SASE spectrum measured by blocking the seed laser (red) and simulation the SASE spectrum after 20 m of NISUS structure (green). The average spacing between spikes in the SASE spectrum is used to estimate the pulse length.



Li-Hua Yu et al.  
Phys. Rev. Lett. 91, 074801 (2003)

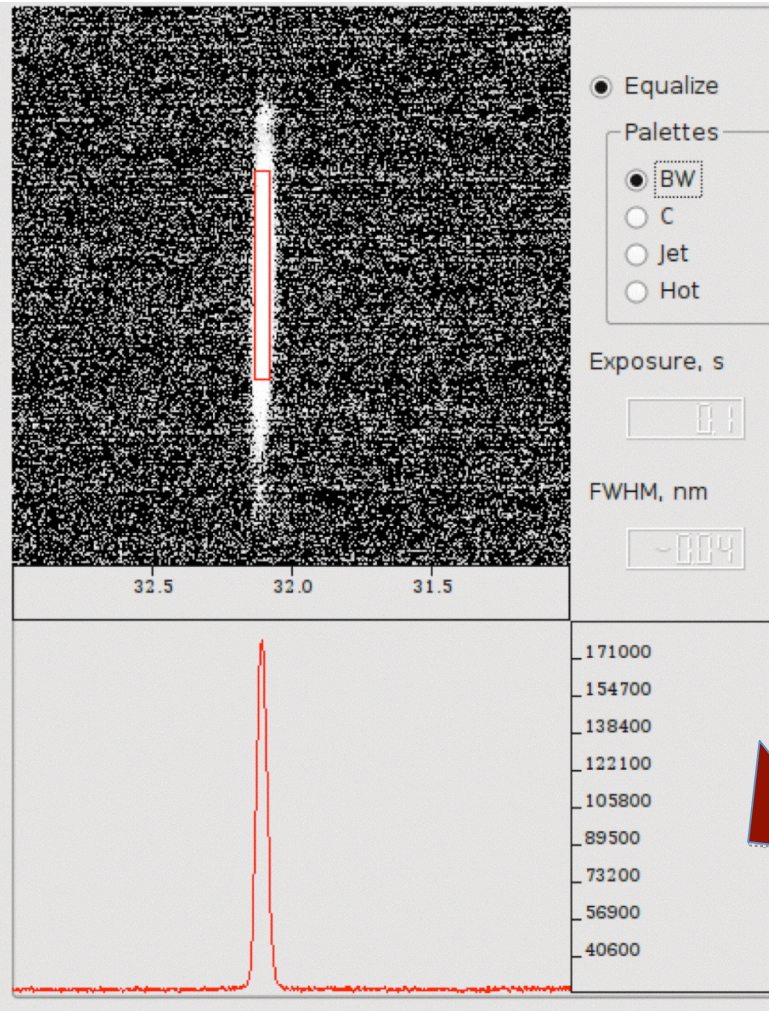
FERMI@Elettra



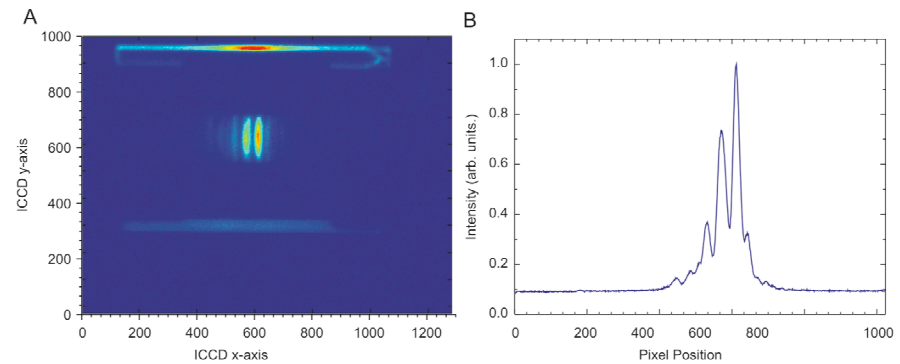
# SASE AND SEEDED

First direct comparison between SASE AND SEEDED in the EUV-Soft X-ray region

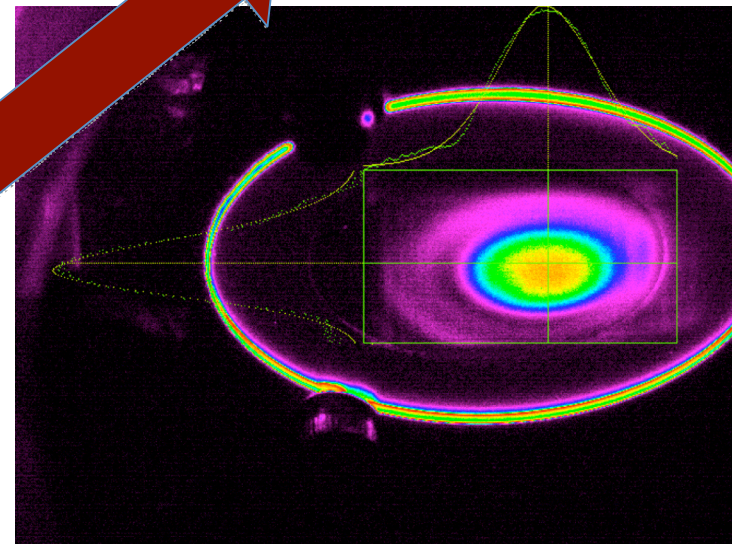
FERMI seeded FEL at 32.2 nm



D. Cocco, C. Svetina, M. Zangrando W. Fawley and E. Allaria

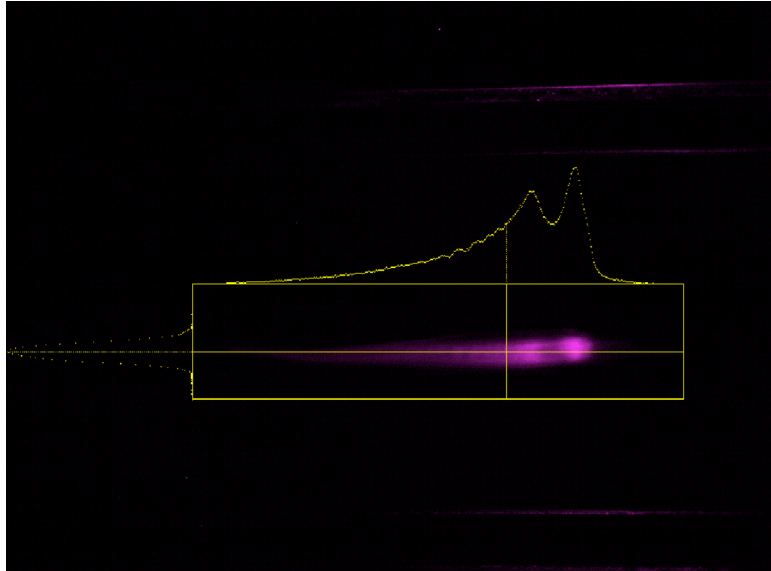


First results from the online variable line spacing grating spectrometer at FLASH, G. Brenner et al. NIMA 2010



W. Fawley, B. Mahieu, E. Allaria

# SEEDING AND BANDWIDTH



Typical used seed laser parameters have been:

Pulse length  $\sim 150\text{fs}$  FWHM with a measured bandwidth at 260nm of 0.8nm (15meV)

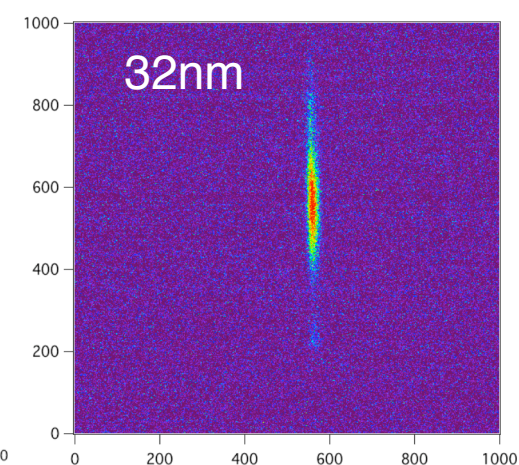
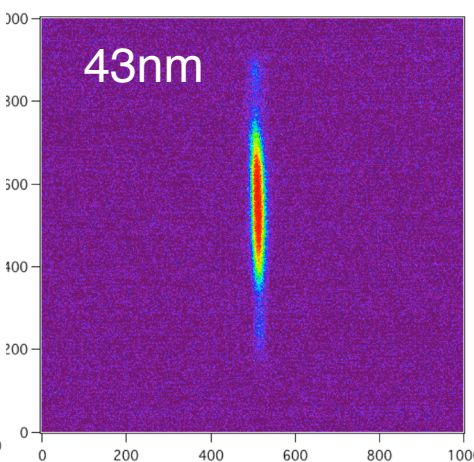
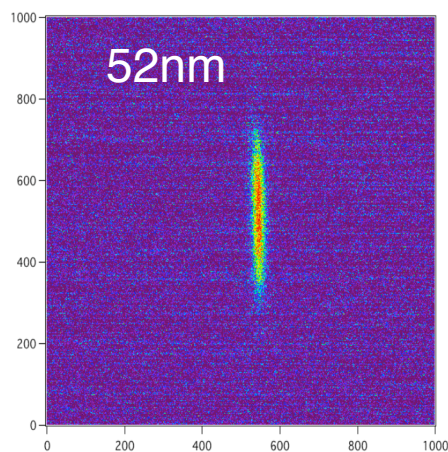
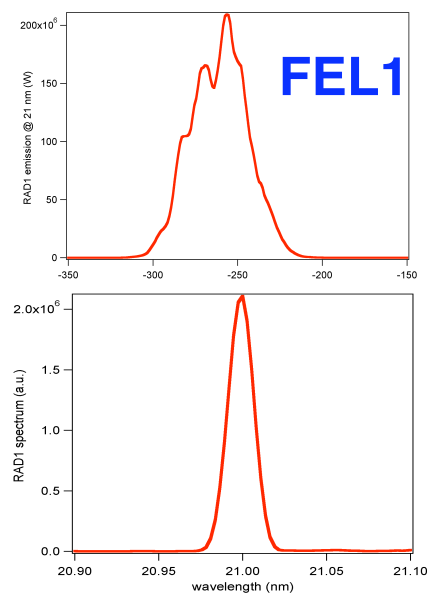


$\lambda$ nm	Number of photons	E $\mu\text{J}$	$\Delta E_{\text{FWHM}}$ meV	$\Delta E_{\text{FWHM}}/E$
52.5	$>10^{13} (*)$	$>40$	30	$1.3 \cdot 10^{-3}$
43.3	$>10^{13} (*)$	$>45$	35	$1.2 \cdot 10^{-3}$
32.5	$>10^{13} (*)$	$>55$	45	$1.2 \cdot 10^{-3}$
20	$4 \cdot 10^{12}$	40	50 100 shots	$8 \cdot 10^{-4}$ 100 shots

(\*) calibration must be confirmed



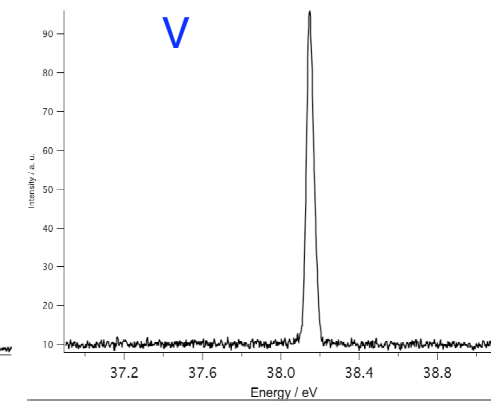
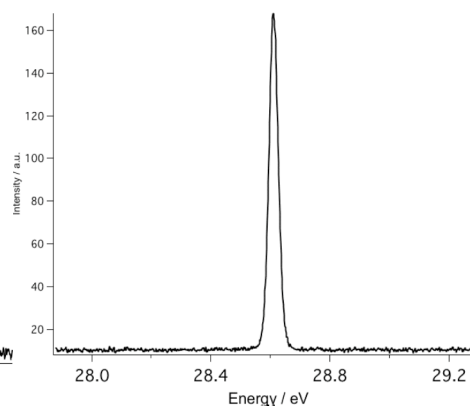
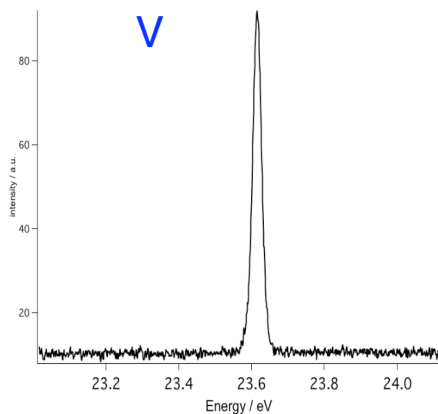
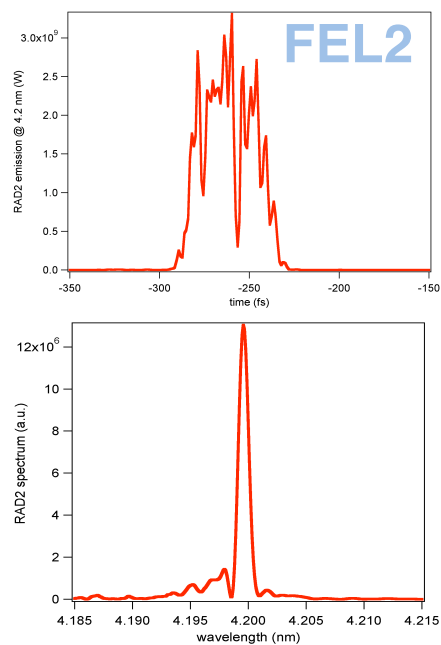
# FERMI FEL1 and FEL2 seeded HGHG layout



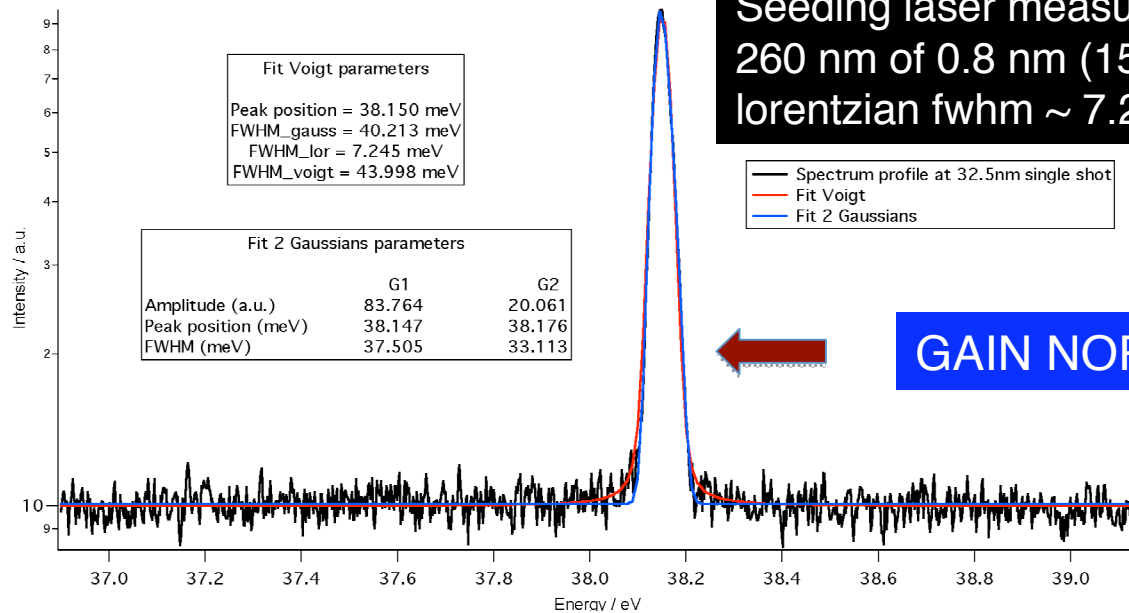
FWHM~29meV  
V

FWHM~33meV

FWHM~44meV  
V



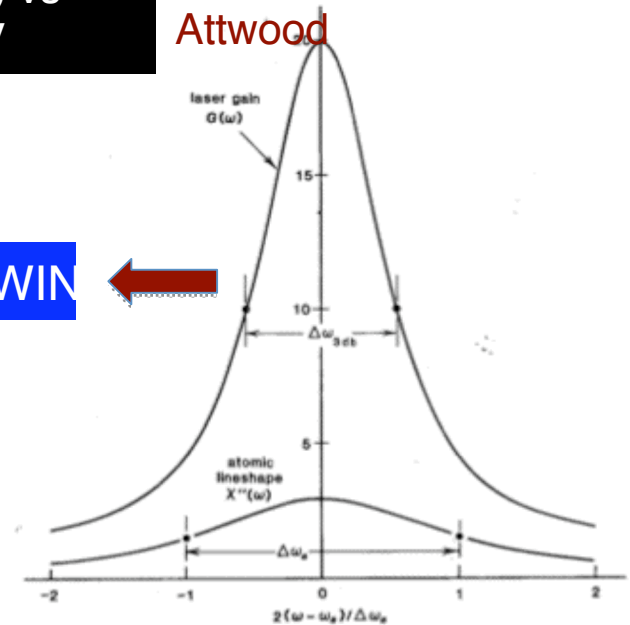
# SPECTRAL LINE ANALYSIS



Seeding laser measured fwhm at 260 nm of 0.8 nm (15meV) vs lorentzian fwhm  $\sim 7.2$  meV

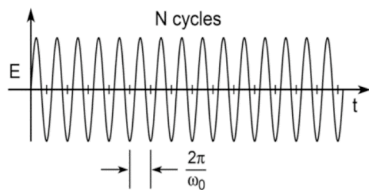
After discussions with Max Zolotarev and Dave Attwood

GAIN NARROWING

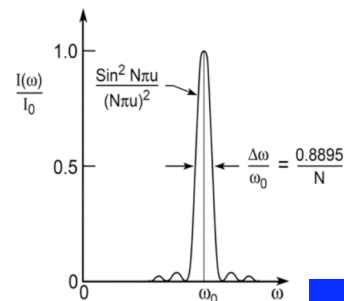


(On-axis radiation,  $\theta = 0$ )

Radiated Wavetrain

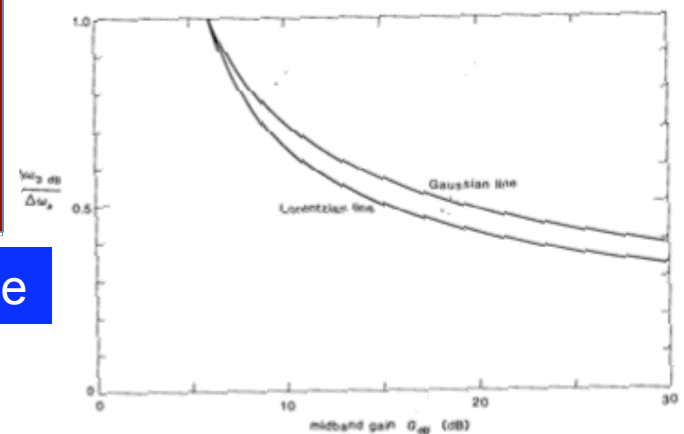


Spectral Distribution



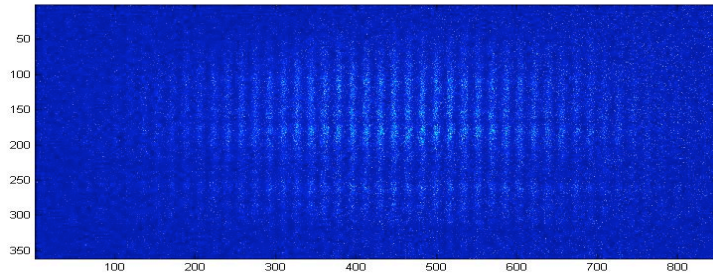
Exponentially amplified e.m. field

Lorentzian line shape

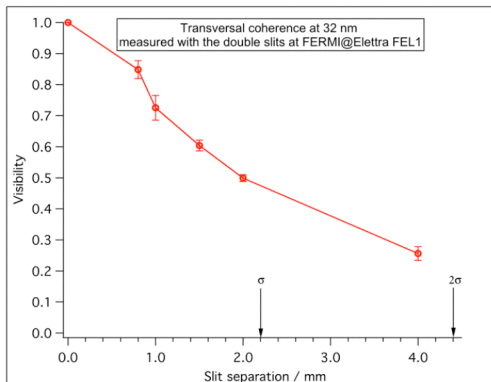




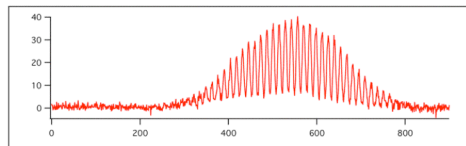
# SASE AND SEEDED COHERENCE



Visibility for a 32nm photon beam (FEL1) measured with the double slits system

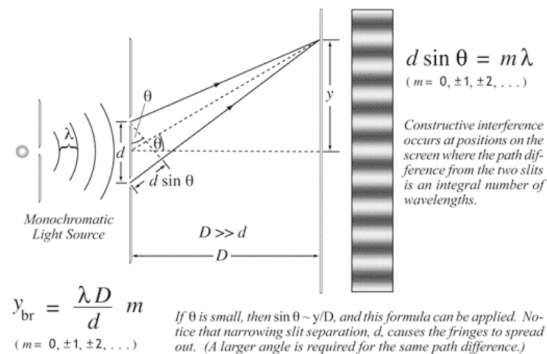


Example of diffraction pattern obtained during the measurements (d=1mm w=40um)



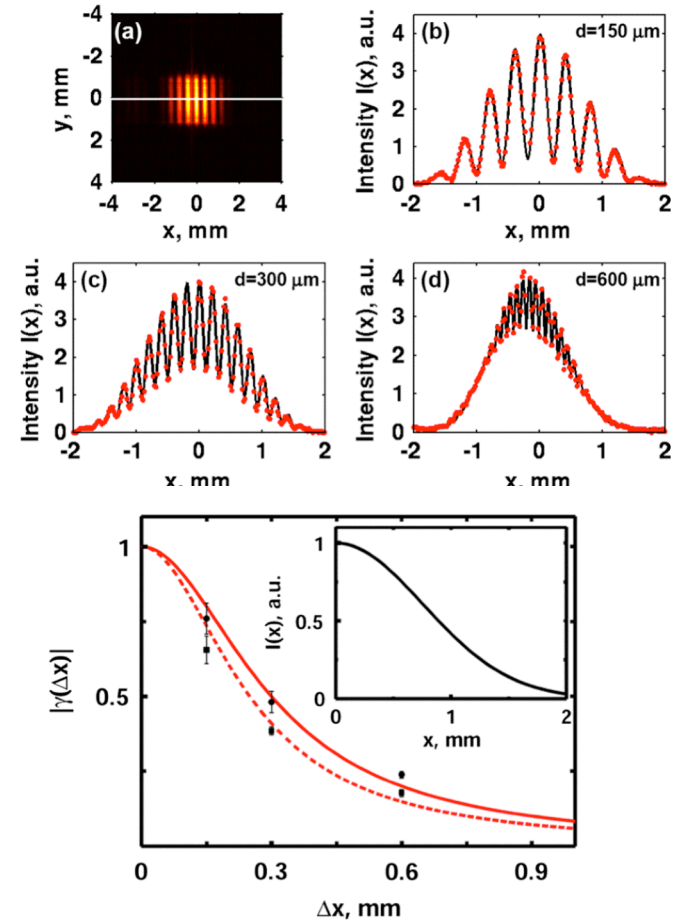
Physics • WAVE OPTICS • XXII • Interference and Diffraction [438] ©

## Young's Double Slit Interference



$$\Gamma(\mathbf{r}_1, \mathbf{r}_2, \tau) = \langle E(\mathbf{r}_1, t) E^*(\mathbf{r}_2, t + \tau) \rangle$$

$$I(P) = I_1(P) + I_2(P) + 2\sqrt{I_1(P)I_2(P)}|\gamma_{12}| \times \cos[\omega\tau - \alpha_{12}(\tau)],$$

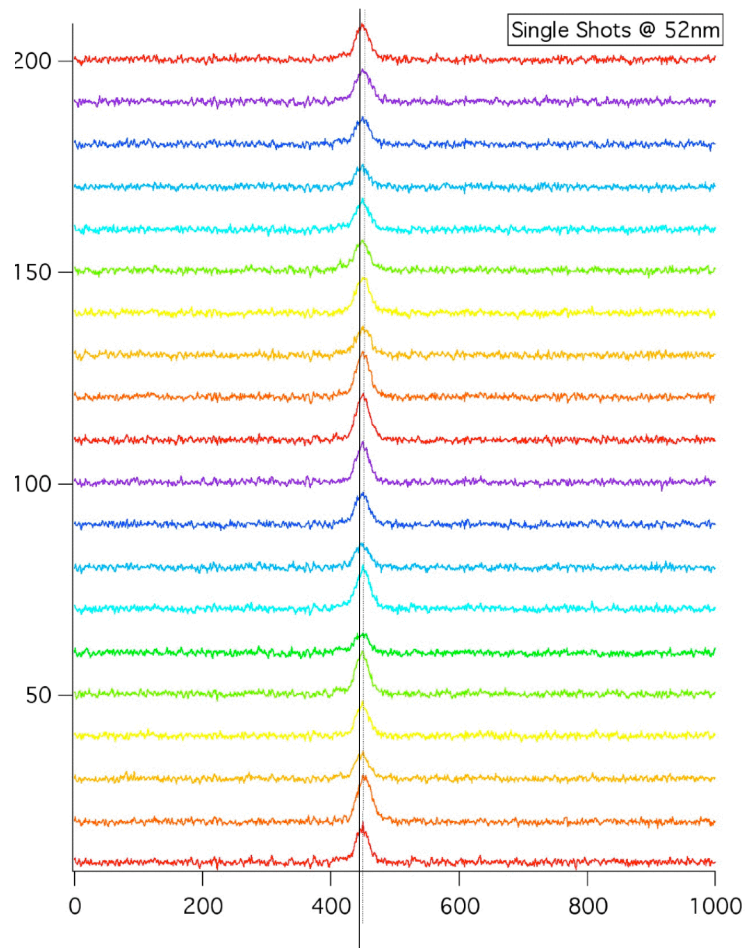


SASE FEL(FLASH) at a wavelength of 13.2 nm  
[see PRL 101,254801(2008)]

FERMI seeded FEL at 32 nm

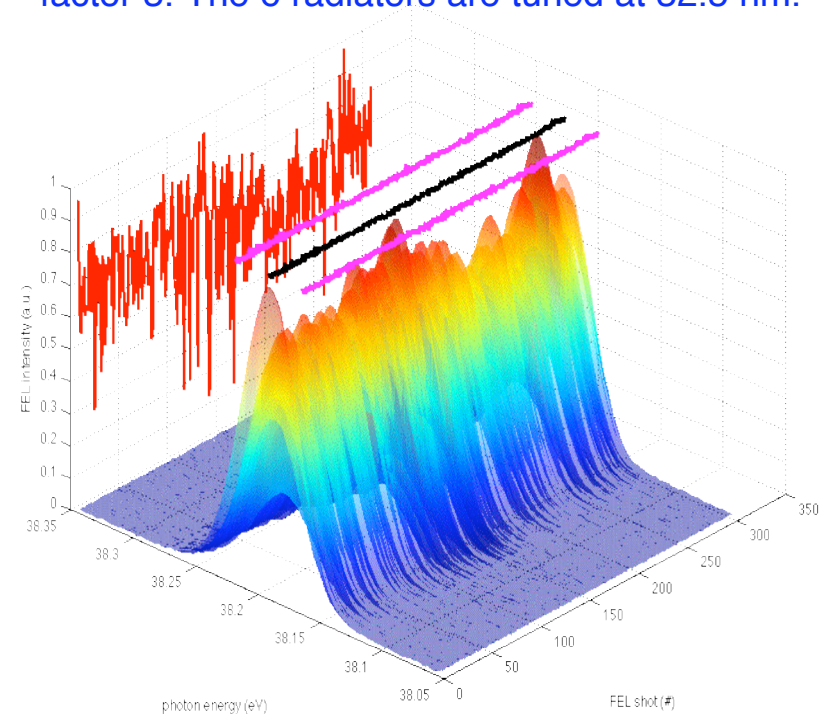
# SASE AND SEEDED BANDWIDTH AND SPECTRAL STABILITY

In addition to the very narrow spectrum FERMI is characterized by a very good spectral stability. Both short and long terms measurements show that the spectral peak move less than  $10^{-4}$ .



D. Cocco, C. Svetina, M. Zangrando

Reported data refer to an electron beam of 350 pC at 1.24 GeV compressed about a factor 3. The 6 radiators are tuned at 32.5 nm.



FEL photon energy  $\sim 38.19\text{eV}$   
Photon energy fluctuations = 1.1meV (RMS)  
FEL bandwidth = 22.5meV (RMS)  
=  $5.9 \times 10^{-4}$  (RMS)  
FEL bandwidth fluctuations = 3% (RMS)

E. Allaria, W. Fawley

# FERMI, measured photon flux

---

## December 2010 (250pC):

Compressed e-beam 43nm with photodiode:	~6 nJ;~1*10 <sup>9</sup>
Compressed e-beam 43nm with spectrometer:	~3 nJ;~5*10 <sup>8</sup>
Compressed e-beam down to 17 nm:	clear evidence of coherent signal

## March 2011 (250pC):

Uncompressed e-beam 65 nm with the DESY gas detector:	~0.3μJ, ~1*10 <sup>11</sup> , ~2MW
Compressed e-beam 65 nm with the calibrated FERMI gas detector:	~3 μJ , ~1*10 <sup>12</sup> , ~20MW

## April 2011 (350pC):

Compressed e-beam 65 nm with spectrometer (average):	~2 μJ, ~6*10 <sup>11</sup> , ~12MW
Compressed e-beam 43 nm with spectrometer (average):	~5 μJ, ~1*10 <sup>12</sup> , ~30MW
Down to ~24 nm	~0.3 μJ, ~4*10 <sup>10</sup> , ~2MW

### **Experimental stations:**

<i>Timex 65 nm with Al filter (March 25<sup>th</sup>):</i>	<i>~135nJ (peak), 70-80nJ (average).</i>
<i>LDM 65 and 52 nm:</i>	<i>estimated from PADReS measurements</i>

## June-July 2011:

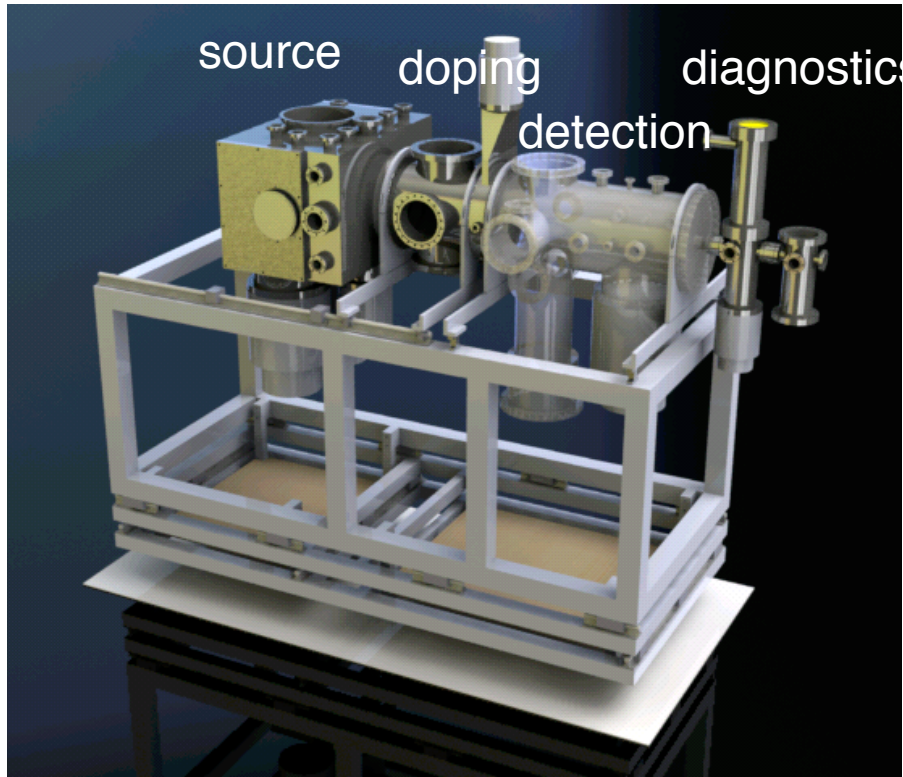
### **350pC-450pC**

Compressed e-beam 52nm with calibrated* photodiode:	~30μJ,	>1*10 <sup>13</sup>	[VERY PRELIMINARY]
Compressed e-beam 43nm with calibrated* photodiode:	~45μJ,	>1*10 <sup>13</sup>	[VERY PRELIMINARY]
Compressed e-beam 32.5nm with calibrated* photodiode:	~100μJ,	>1*10 <sup>13</sup>	[VERY PRELIMINARY]
Compressed e-beam 20nm with calibrated* photodiode:	~40μJ,	~1*10 <sup>12</sup>	[VERY PRELIMINARY]

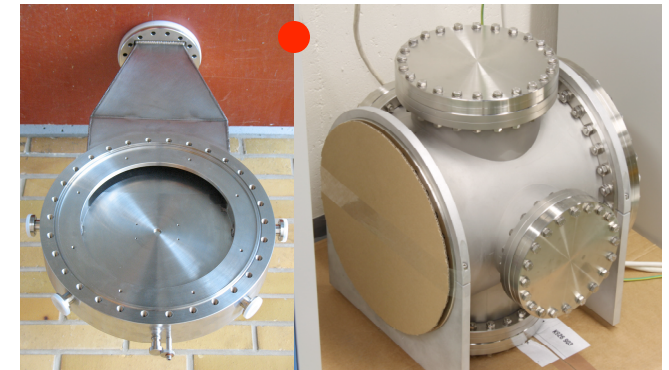
FERMI CT: E. Allaria, M. Trovo', P. Craievich, S. Di Mitri, G. Penco\* **Photodiode calibration to be confirmed**



# LDM end-station



- FRAME CONSTRUCTION COMPLETED
- SOURCE CHAMBER COMPLETED
- DOPING CHAMBER COMPLETED
- DIAGNOSTIC CHAMBER COMPLETED
- **DETECTION CHAMBER BEING FINALIZED**



C. Callegari and LDM TEAM



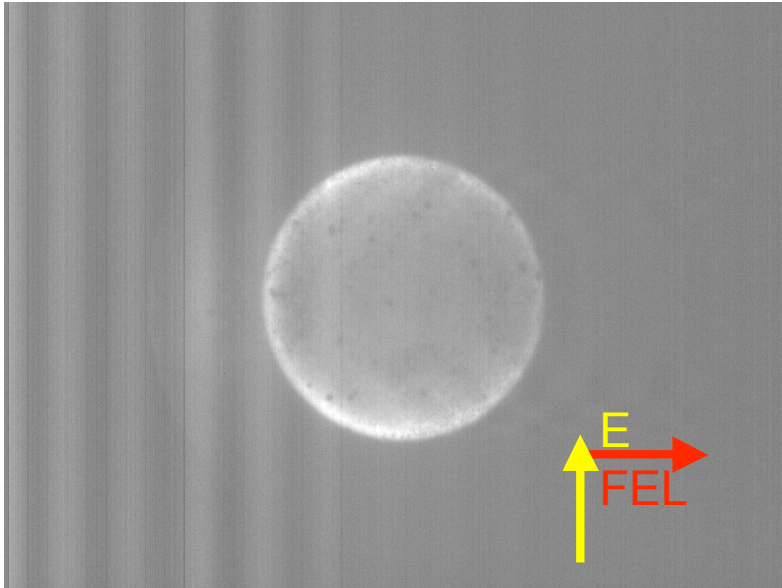
## LDM First experiments

First FERMI polarization data:

He photoelectron spectra

(averaged)

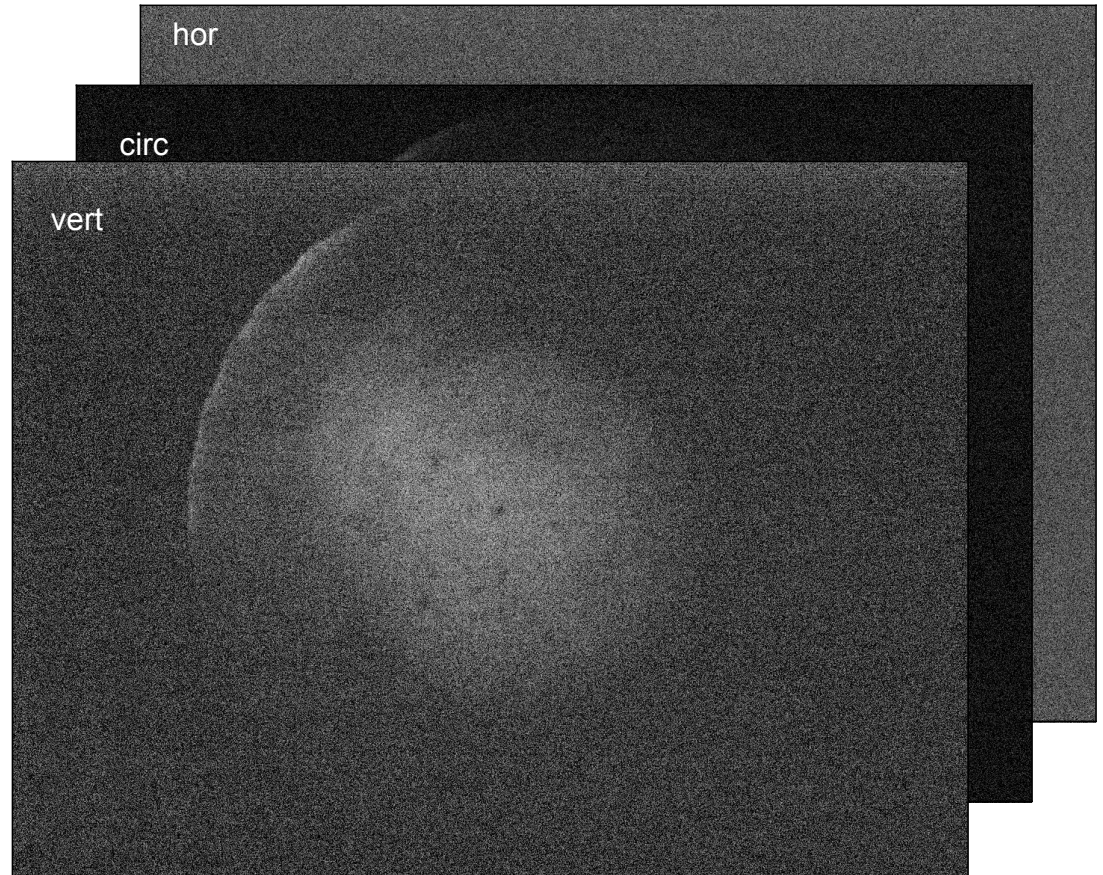
$h\nu=43$  nm



First FERMI spectrum:

Ar photoelectron spectrum

$h\nu=65$  nm, horizontal polarization



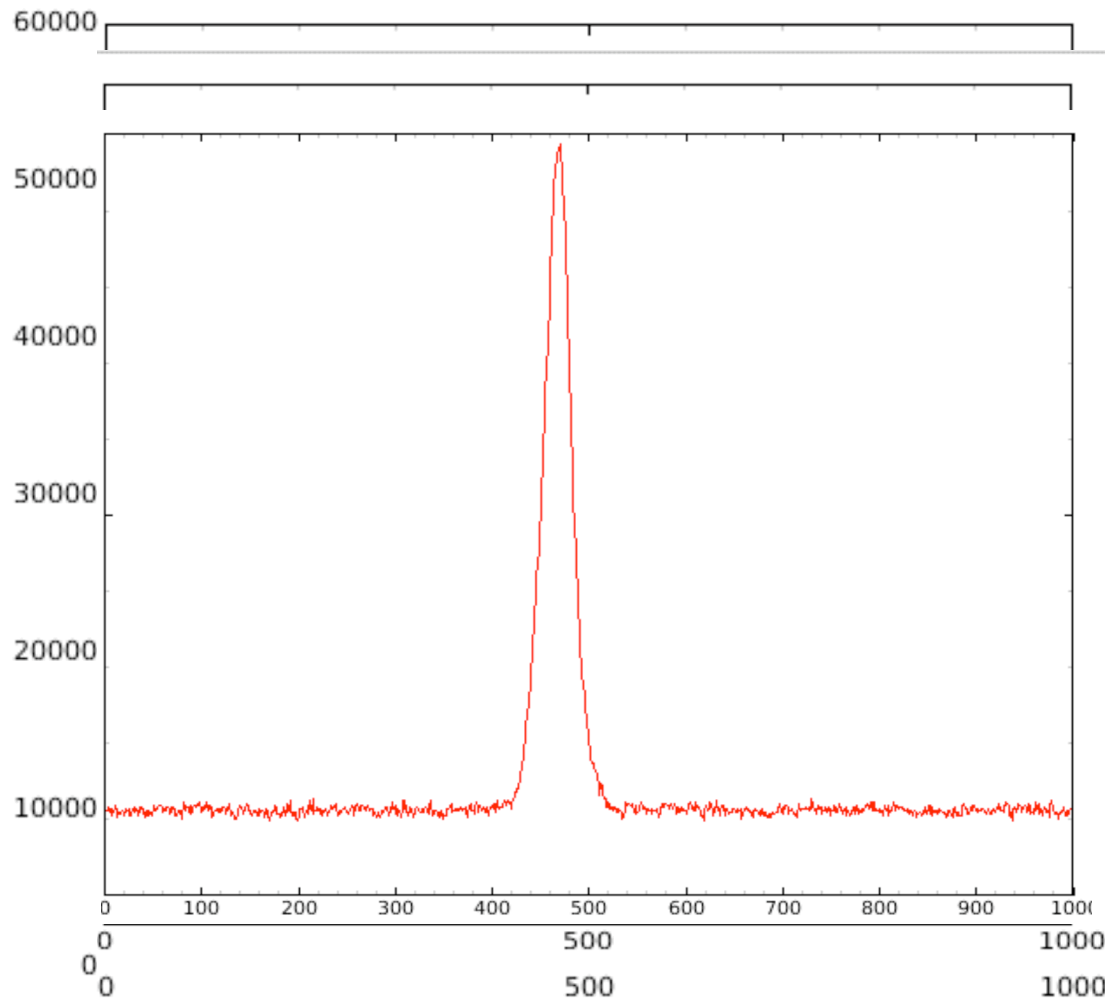
### INSTITUTIONS:

Sincrotrone Trieste; Technical University Berlin (TUB); University of Freiburg (UF); University of Milan; University of Rome

\*EXP HALL: Michele Alagia (CNR-IOM), Paola Bolognesi (CNR-IMIP), Carlo Callegari, Marcello Coreno (CNR-IMIP), Monica de Simone (CNR-IOM), Michele Devetta (UniMI), Vitaly Feyer, Raphael Katzky (UF), Antti Kivimaki (CNR-IOM), Victor Lyamayev (UF), Patrick O'Keeffe (CNR-IMIP), Yevheniy Ovcharenko (TUB), Kevin Prince, Robert Richter, Rudi Sergo, Stefano Stranges (UniRM).

\*CONTROL ROOM: Carlo Spezzani.

# FEL tuning



A small FEL tuning around 52nm has been achieved by changing the seed laser wavelength of 1 nm (0.4%).

Following the seed laser wavelength tuning the undulator resonance has been changed accordingly to maximize the FEL power.

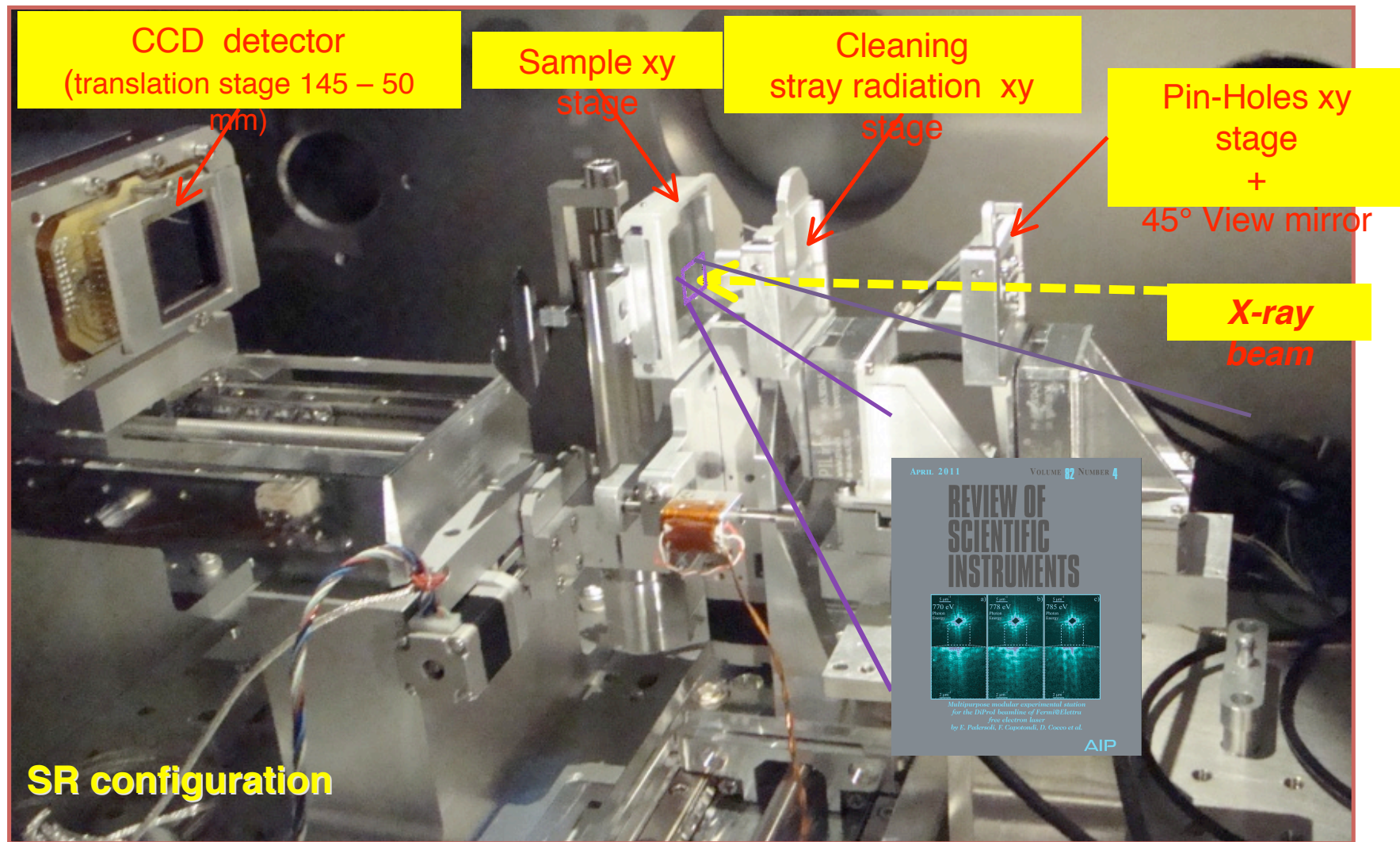
$\Delta\lambda_{\text{seed}}/\lambda_{\text{seed}} = 0.4\%$ ,  
undulators gap tuning  
around 52 nm  
→ **on-line tunable  $\lambda_{\text{FEL}}$**

These data collected by: E. Allaria, B. Mahieu, W. Fawley, B. Diviacco, M. Musardo, L. Froehlich, C. Svetina, C. Spezzani, M. Danailov, S. Demidovich, C. Callegari, M. Zangrando



# DiProl: test on Nanospectroscopy

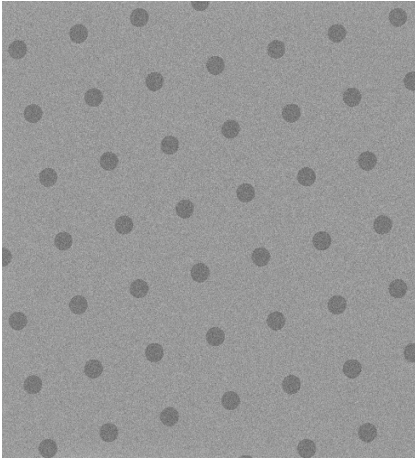
Detector Configuration used at Nanospectroscopy is in direct detection mode as should be provided with proper detector for FERMI-2. **Demonstrated feasibility of resonant magnetic imaging using circular polarisation .**



M. Kiskinova and DiProl Team

# First use of FEL light in experimental chambers

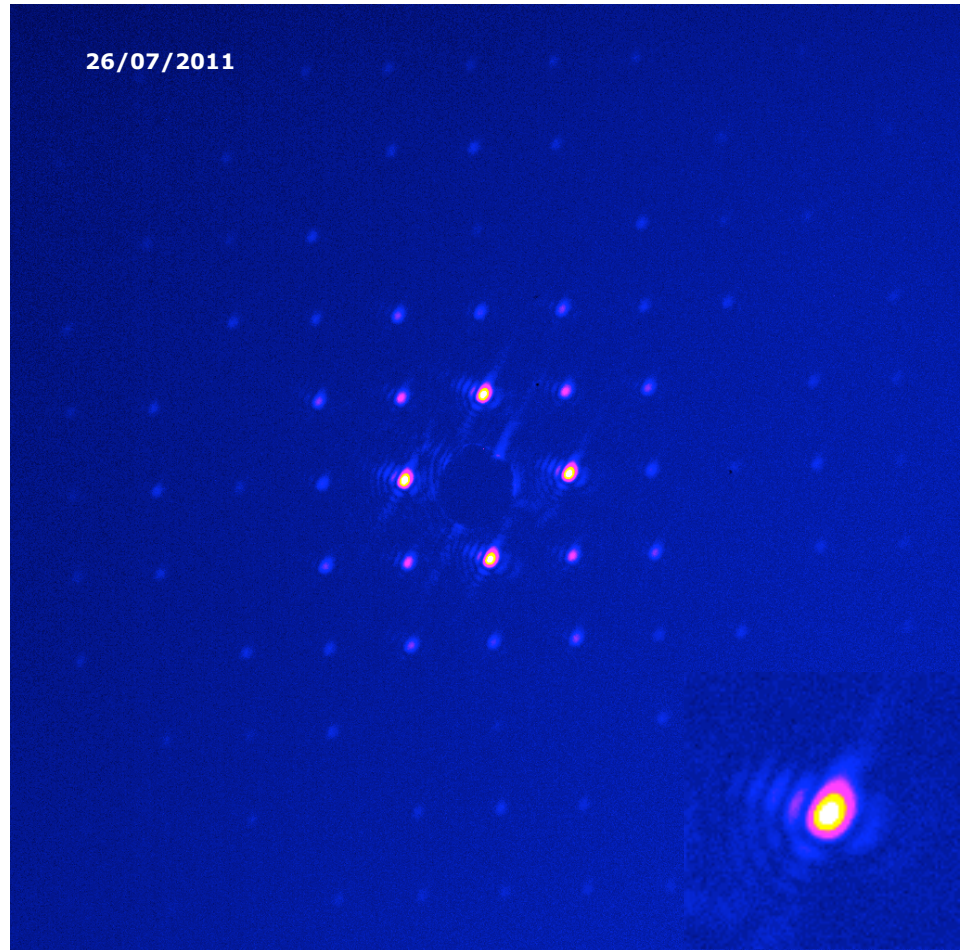
---



FEL radiation at 32.5 has been used to start with the activities on the DiProl experimental chamber.

Due to an issue on the focusing mirror the small FEL spot has been produced by means of a 20  $\mu\text{m}$  pin-hole from QuantiFoil. FEL signal has been also filtered by using two Al filters (400 and 800 nm respectively)

In these conditions it has been possible to record coherent scattering images from a periodic array by integrating the CCD signal over few minutes.



F. Capotondi, E. Pedersoli, R. Menk, M. Kiskinova and  
H. Chapman et al. (CFEL-DESY), J. Hajdu et al. (Uppsala), M. Bogan et al. (SLAC), M. Pivovarov, A. Nelson et al. (LLNL)  
S. Spampinati, S. Bassanese, E. Allaria



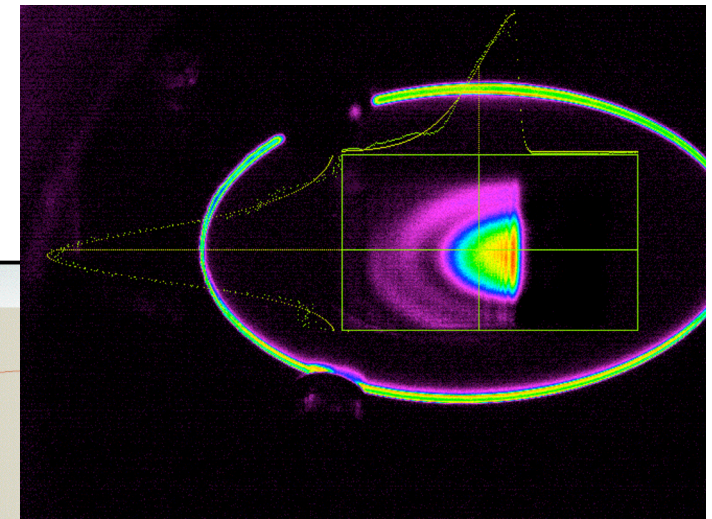
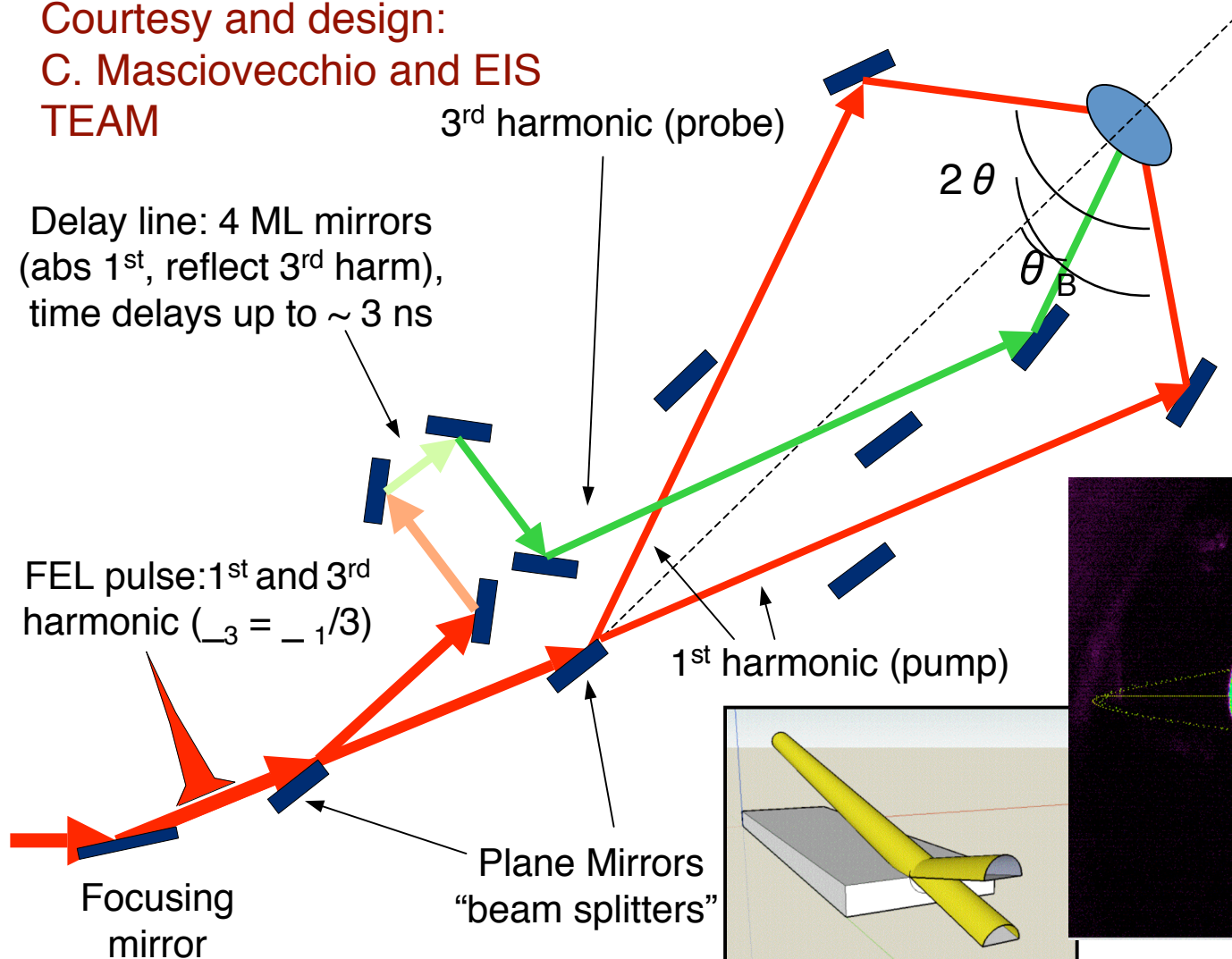
# TIMER: optical layout



Courtesy and design:  
C. Masciovecchio and EIS  
TEAM

Delay line: 4 ML mirrors  
(abs 1<sup>st</sup>, reflect 3<sup>rd</sup> harm),  
time delays up to  $\sim 3$  ns

FEL pulse: 1<sup>st</sup> and 3<sup>rd</sup>  
harmonic ( $\omega_3 = \omega_1/3$ )



W Fawley

*R. Cucini et al et al., NIMA (2011)*

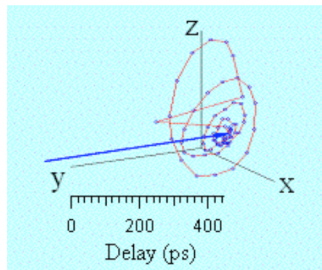




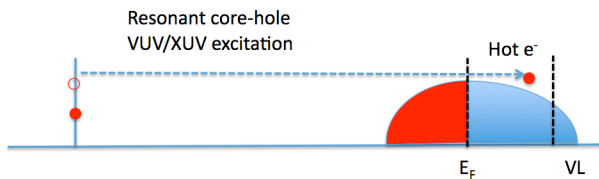
# Magnedyn scientific proposals and cases

## Ultrafast demagnetization

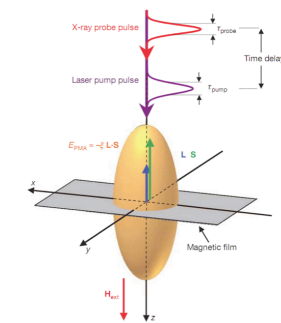
Ultrafast demagnetization after optical excitation



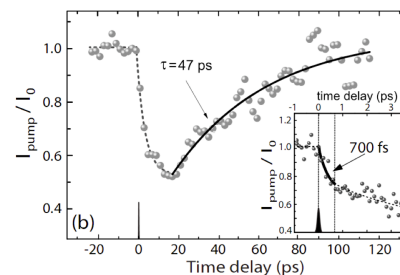
## Ultrafast demagnetization after core hole excitation



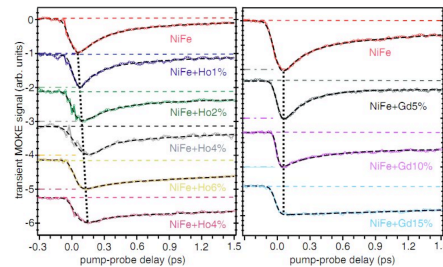
## temporal evolution of the orbital moment anisotropy



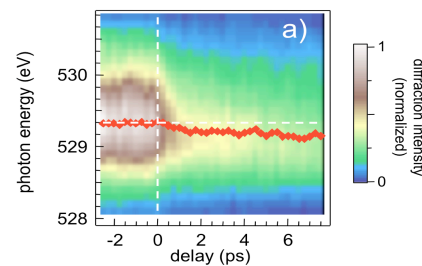
## magnetization dynamics in complex materials



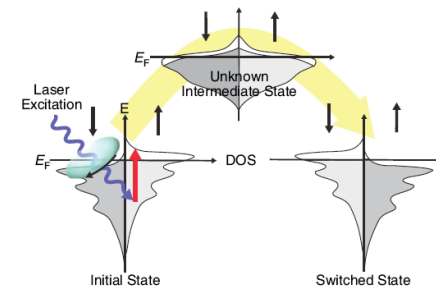
## Magnetic dynamics in dilute systems.



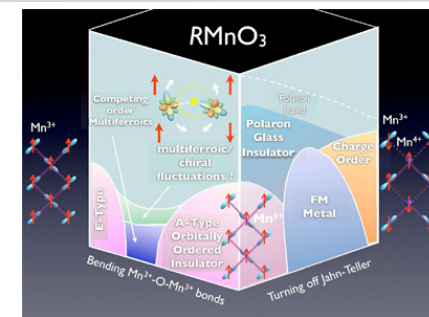
## Transient phases in complex materials.



## Phase Transitions in Advanced Magnetic Materials.



## Single shot experiments And magnetic fluctuations





## Future Scenario

- The way to produce fully coherent X-ray radiation is paved. If also the second stage for the HGHG will be proved a new technology will be available.
- Tunability
- Variable polarization
- Full coherence
- High repetition rate

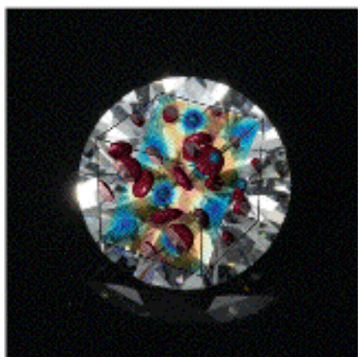
The future scenario

Coherent X-ray Optics

Quantum X-ray optics

Stroboscopic phase tomography

An extraordinary effort is needed to develop a suitable science program



XPDC imaging. Tamasaku et al.<sup>4</sup> use their parametric down-conversion-based technique to investigate the response of diamond to ultraviolet light at a resolution as small as 0.54 Å.

COURTESY OF KENJI TAMASAKU

### NONLINEAR X-RAY OPTICS

## The next phase for X-rays

Phase information can be obtained from inelastically scattered X-rays by combining parametric down-conversion with tunable quantum interference. This is a step towards putting this nonlinear phenomenon to a practical use in the X-ray regime: investigating the optical response of chemical bonds at their electron-volt and subnanometre scales.

Bernhard Adams

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news & views



## Future road map

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### FEL-1

FEL-1 optimization is expected to be concluded in 2011 and the users' program started in 2012

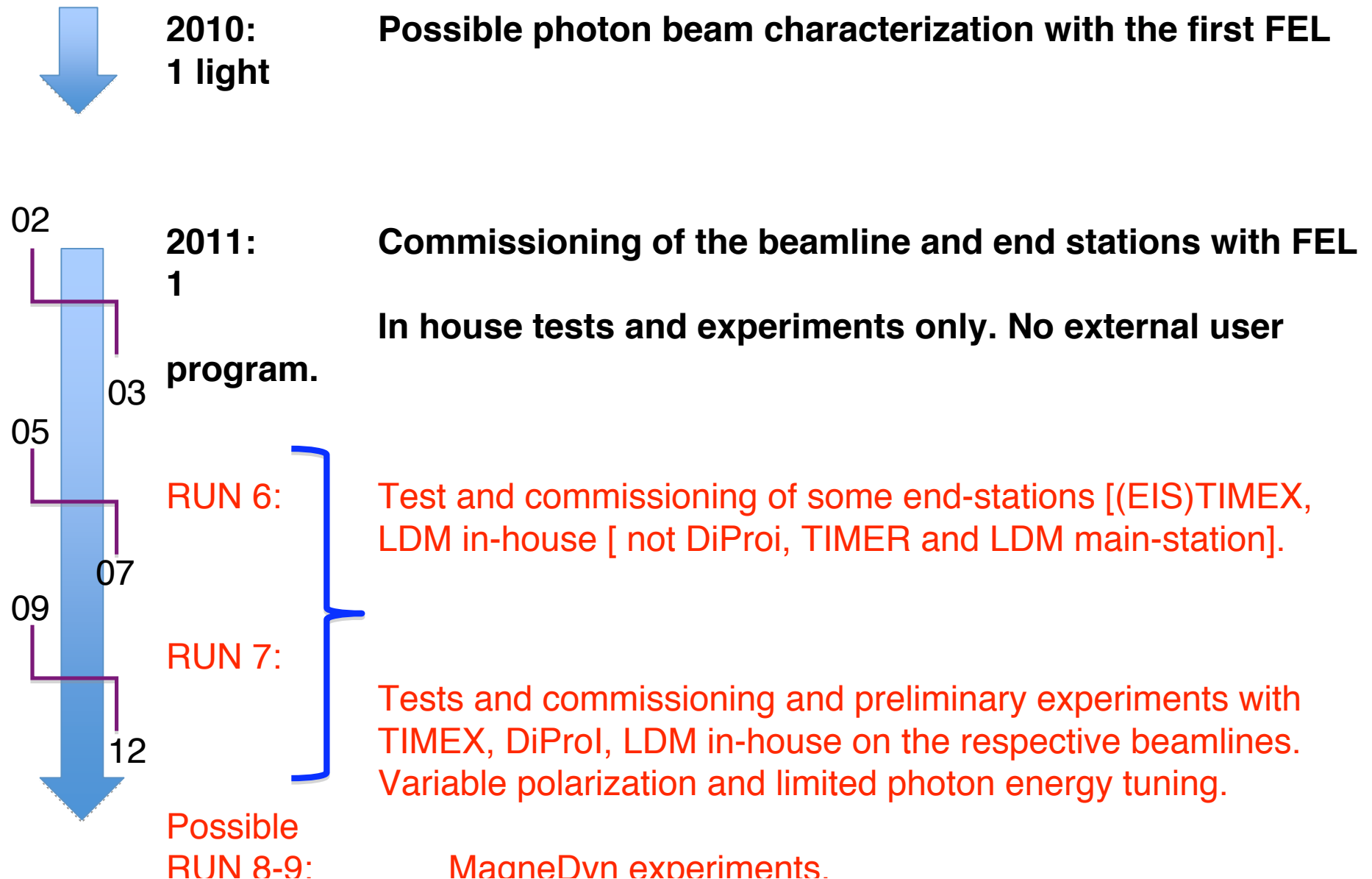
Two projects have been already founded for implementing HHG sources to be tested as a seed on FERMI in the 30 nm range and more long term below 10nm.

When FEL-2 become online FEL-1 could be temporary configured for HHG tests (*end 2012?*).

### FEL-2

Prove HHG with FEL2 (major goal in 2012)

FEL-2 has been already shown to be almost compatible to ECHO scheme. A possible temporary modification could be done in agreement with the users for testing ECHO at 50<sup>th</sup>



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