

Beam dynamics considerations for a high repetition rate photoinjector

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CBP SEMINAR
NOVEMBER 16 2012

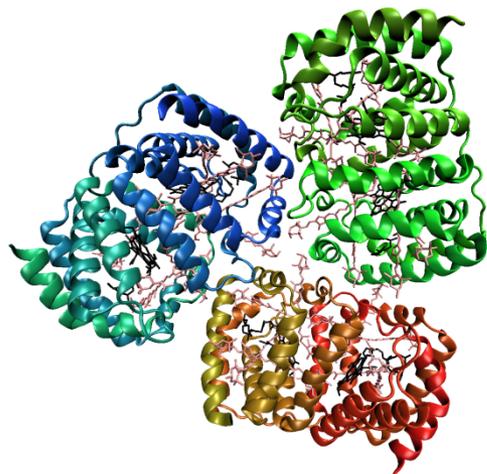
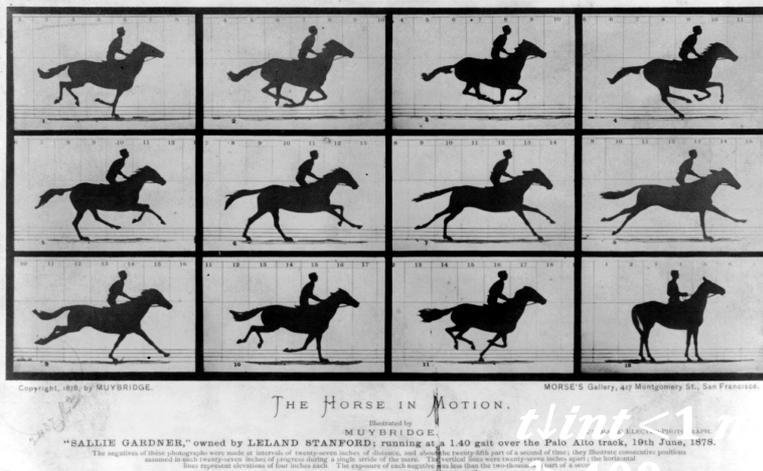
1. Introduction and Motivation
2. Design and Optimization of the NGLS Injector
3. Beam dynamics for the Advanced Photoinjector
EXperiment
4. Emittance diagnostics based on OTR

Shorter length, smaller size, brighter, coherent light

Le derby d'Epsom, painting by Théodore Géricault, 1821



The horse in motion by Eadweard Muybridge, commissioned by Leland Stanford, 1878



?

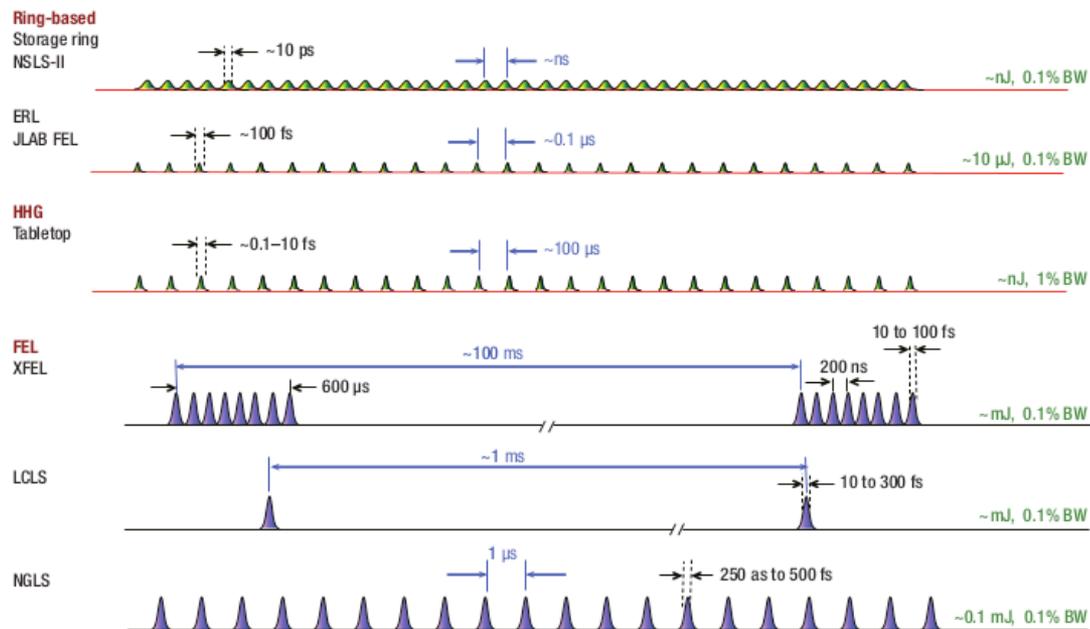
The peridinin-chlorophyll-protein light-harvesting complex.

Sources

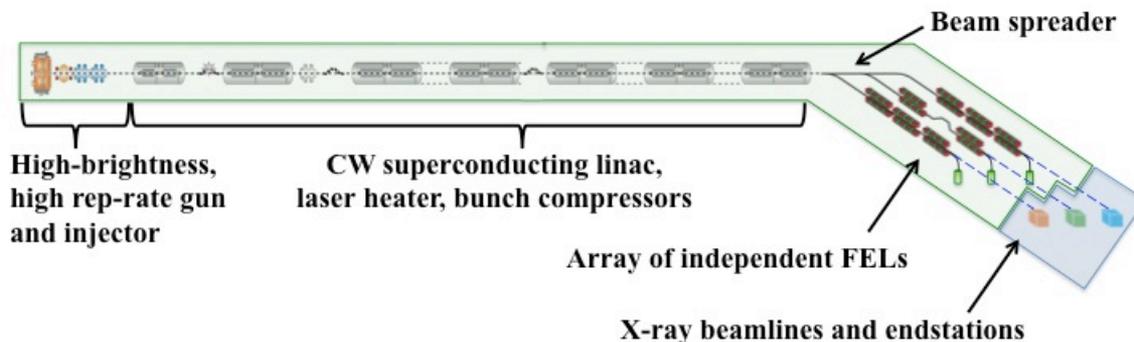
- http://upload.wikimedia.org/wikipedia/commons/f/f3/Jean_Louis_Th%C3%A9odore_G%C3%A9ricault_001.jpg
- http://upload.wikimedia.org/wikipedia/commons/7/73/The_Horse_in_Motion.jpg
- <http://upload.wikimedia.org/wikipedia/commons/1/19/Peridinin-chlorophyll-solenoid-1ppr.png>

NGLS: A high repetition rate soft xray FEL

Linac driven FELs are a proven tool to produce xray pulses with high power, small bandwidth and short length

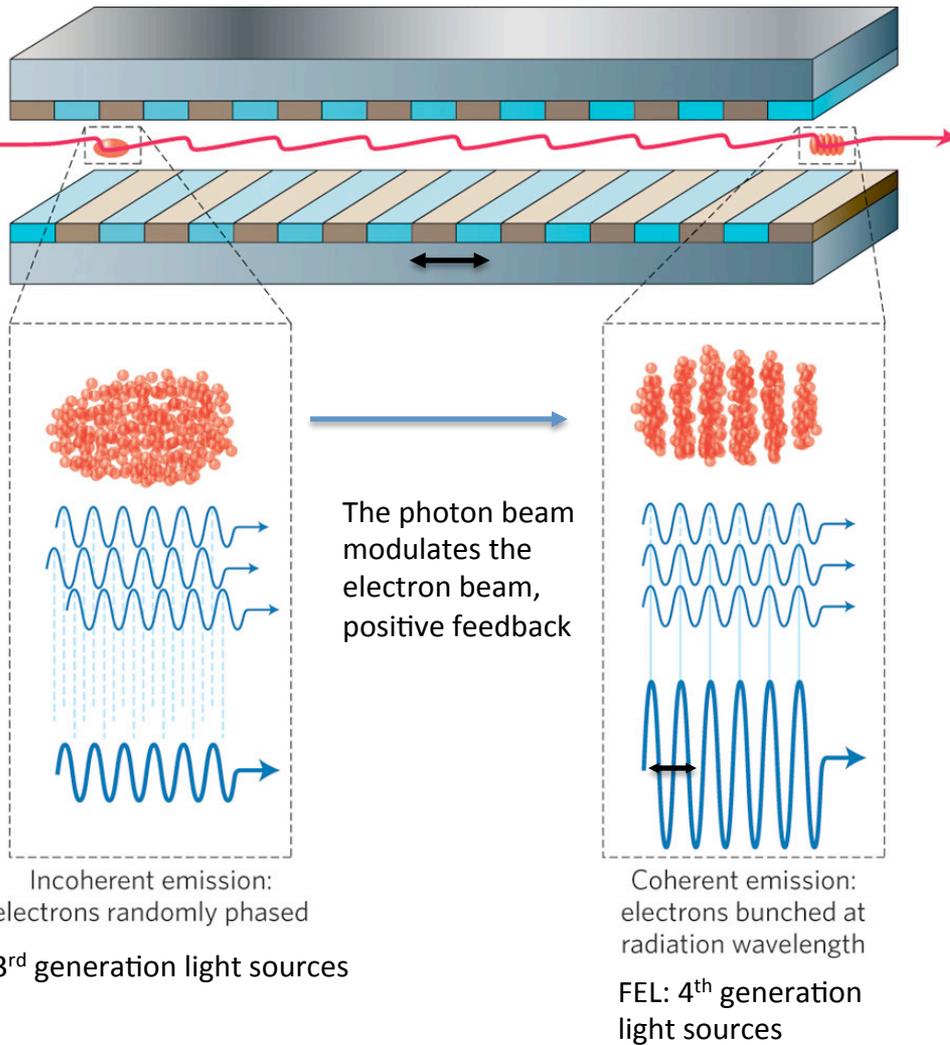


NGLS will bridge the gap between high repetition rate, low peak power storage rings and low rep. rate, high peak power existing FELs



Source: NGLS CDO, CDR

The FEL process



Important FEL parameters

- : radiation wavelength
- : undulator parameter
- : Exponential power gain
- : Gain length (1-d theory)

Short list of requirements for lasing

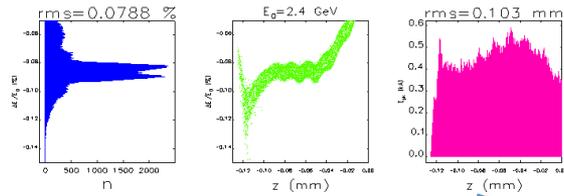
- : Resonance condition
- : Transverse overlap of electrons and photons

high enough for reasonable undulator length

Stringent requirements on transverse and longitudinal beam quality

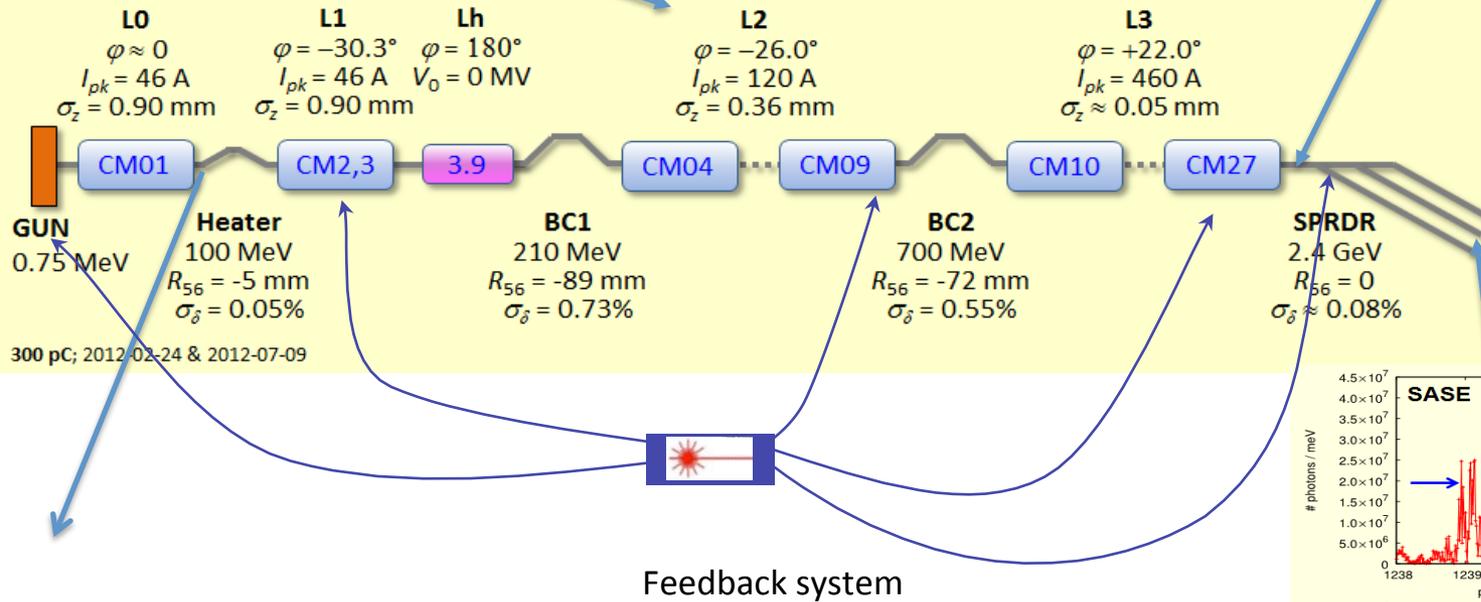
Source: Brian W. J. McNeil & Neil R. Thompson Nature Photonics 4, 814–821 (2010)

NGLS: A single pass system with multiple components



Linac: Accelerate and compress while avoiding trans. emittance degradation and microbunching

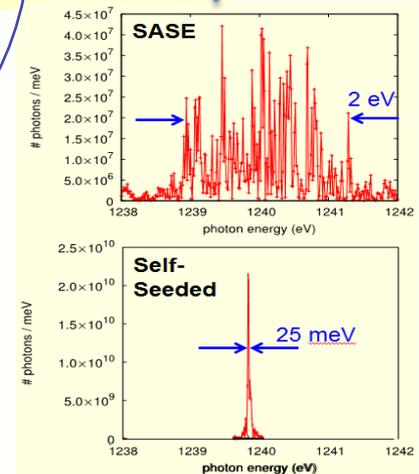
Spreader: Beam distribution to individual beamlines



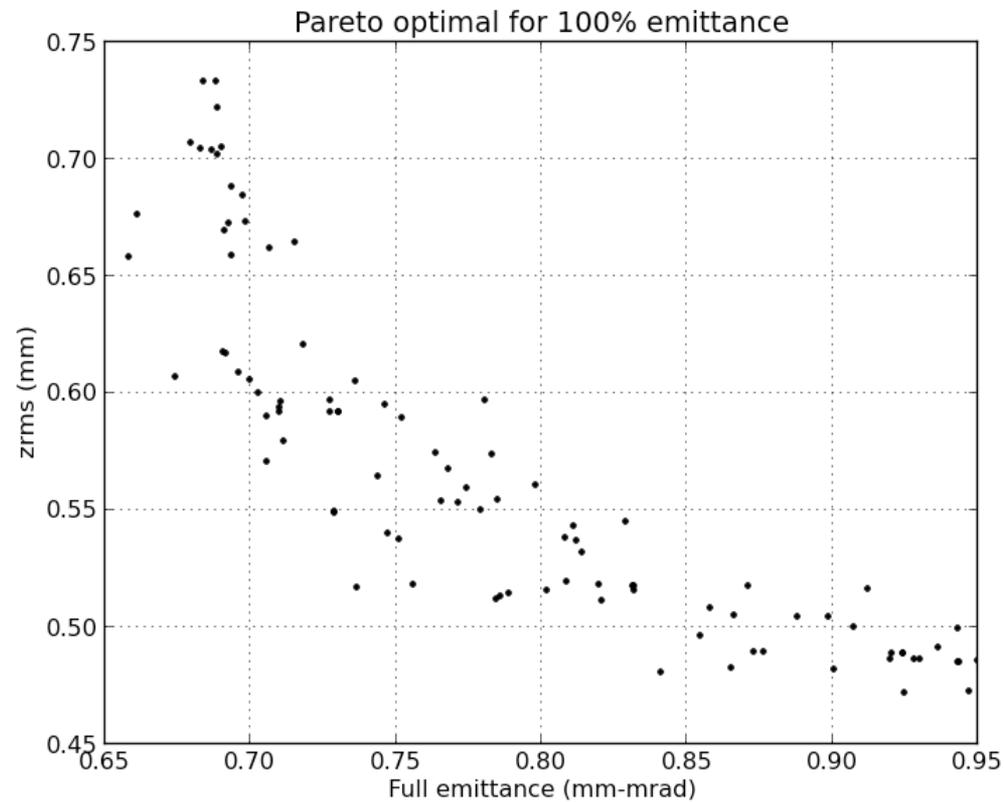
Exit of the NGLS Injector:
The beam quality cannot be improved downstream
(unlike synchrotrons)

Feedback system

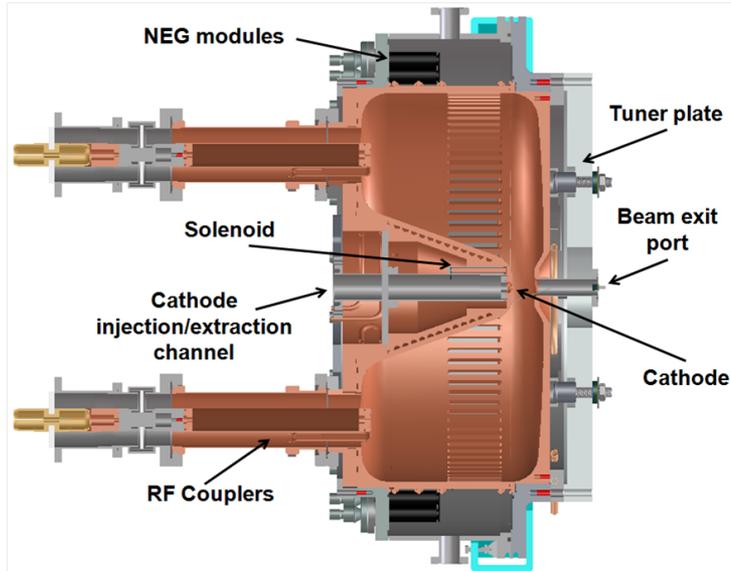
Photon beam after FEL



Injector optimization



Normal Conducting, Continuous Wave RF Gun



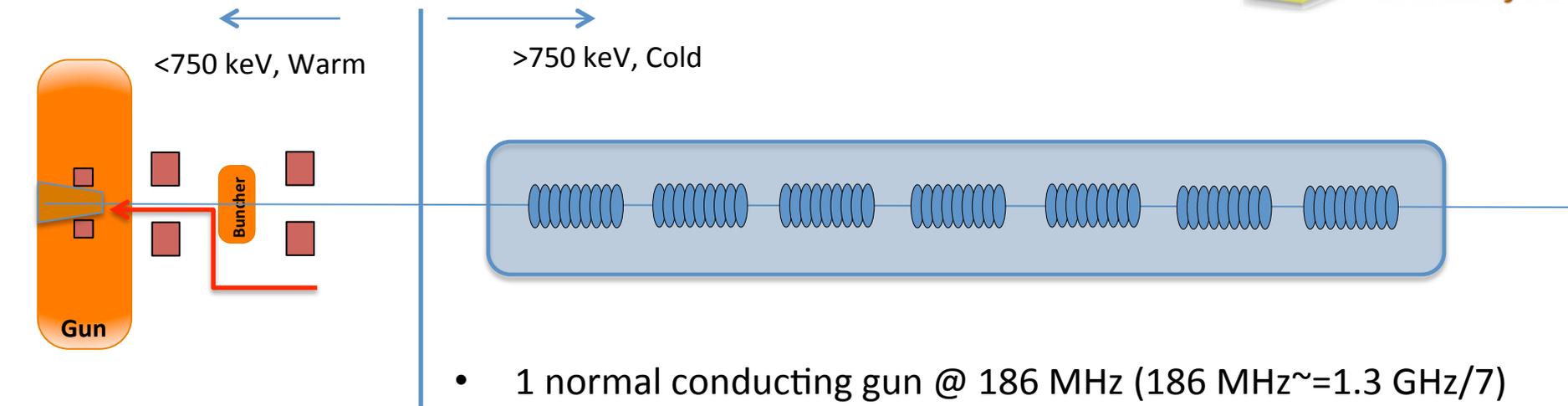
Frequency	187 MHz
Operation mode	CW
Gap voltage	750 kV
Field at the cathode	19.47 MV/m
Q_0	30887
Shunt impedance	6.5 M Ω
RF Power	87.5 kW
Stored energy	2.3 J
Peak surface field	24.1 MV/m
Peak wall power density	25.0 W/cm ²
Accelerating gap	4 cm
Diameter/Length	69.4/35.0 cm
Operating pressure	< 10 ⁻¹¹ Torr

- Based on proven technologies
- Normal conducting RF cavity
- Large size for easy cooling
- Good vacuum to accommodate high QE photocathodes

Gun Technology Fully Demonstrated
The gun is the basis of the APEX project
(more on this later)

Fernando Sannibale, Daniele Filippetto, Christos F. Papadopoulos

NGLS Injector Layout



- 1 normal conducting gun @ 186 MHz (186 MHz \sim 1.3 GHz/7)
- 2 solenoids (emittance compensation)
- 1 single-cell buncher cavity @1.3 GHz (for compression)
- 7 9-cell TESLA cavities @1.3 GHz (velocity bunching + emittance compensation)
- 1st cryomodule is identical to the downstream cryomodules

- Solenoid magnet
- Warm single-cell RF cavity
- ⊞ Cold multi-cell RF cavity
- ↪ Laser pulse
- ▭ Cryomodule

Beam dynamics implications of the NGLS Injector



Compare with what works already

- Longer RF wavelength (186 MHz vs 2.85 GHz for LCLS)
- Lower gradient at the cathode (20 MV/m vs 140 MV/m for LCLS)
- Lower beam energy at the exit (0.75 MeV vs 6 MeV for LCLS)

Space charge is more important in our case

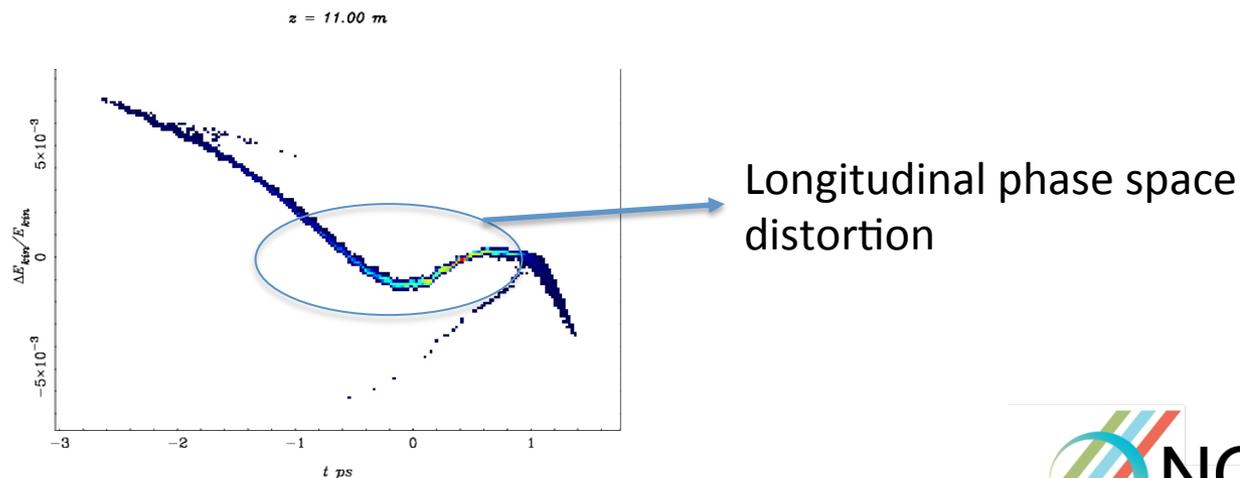
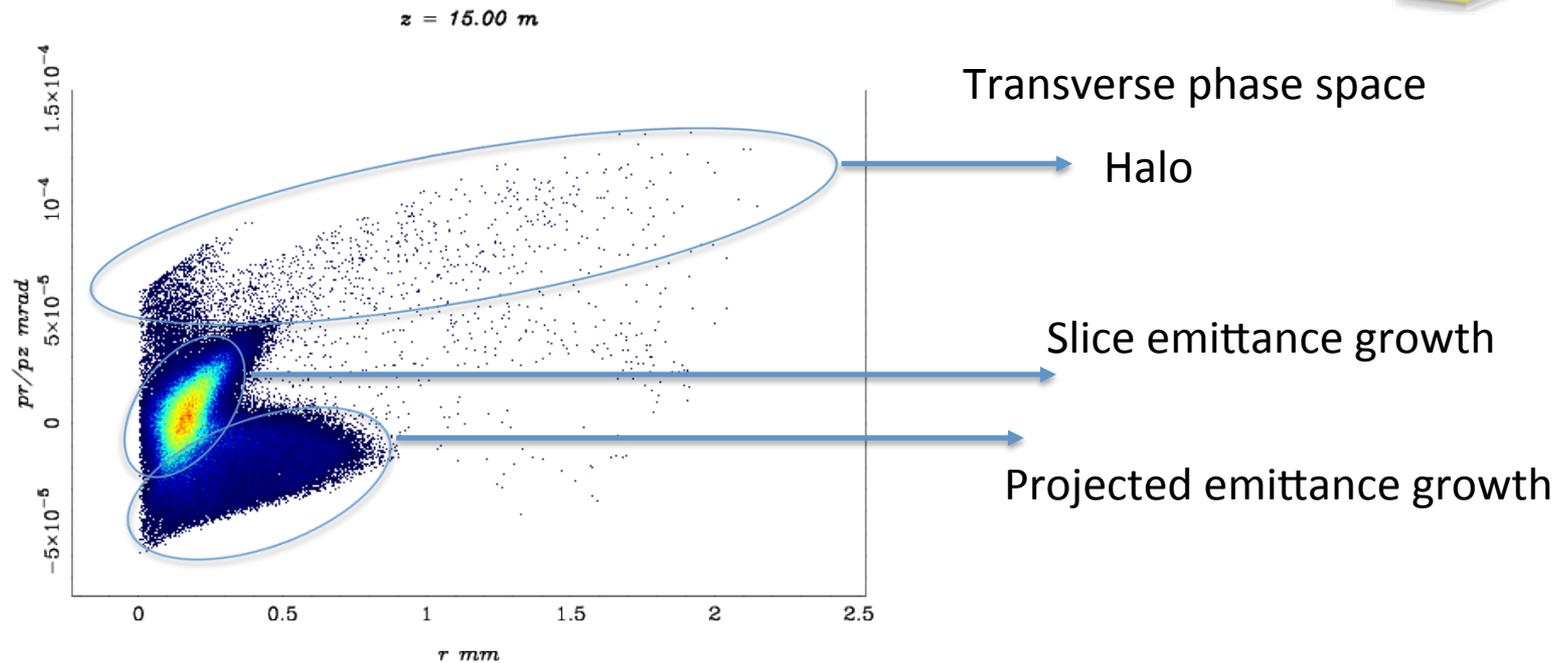
Transverse SC force

Longitudinal SC force

geometric factor,
depends on bunch shape

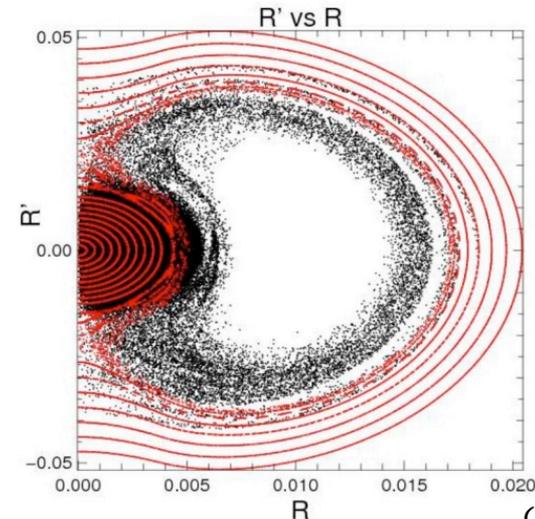
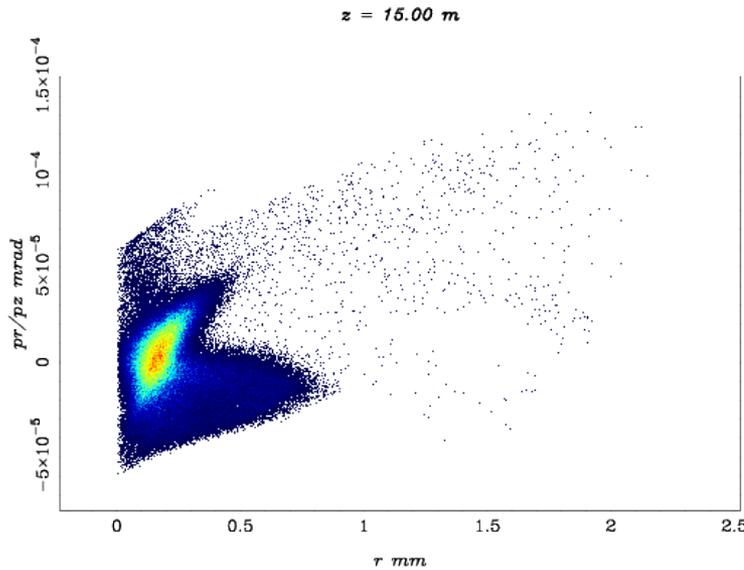
Space charge (linear and nonlinear) can dilute the quality of the phase space

Phase space distortions due to space charge



Beam Halo

Black: WARP simulation of a space charge dominated beam (UMER)
 Red: Analytical particle-core model



$$\frac{d^2 r}{dz^2} + k_0^2 r - F(r) = 0 \quad F(r) = \begin{cases} mv_z^2 \frac{K}{R^2} r & r \leq R(z) \\ mv_z^2 \frac{K}{r} & r > R(z) \end{cases}$$

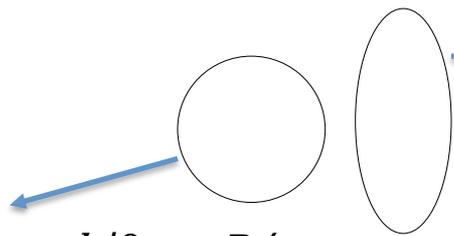
Particles that resonate with the beam envelope oscillations may be kicked to high amplitude oscillations and create beam halo.

We expect most beam losses to be due to dark current (more on this later)

C. F. Papadopoulos, PhD Thesis, 2009
 Gluckstern 1994
 Lagniel 1994
 Wangler 1998

Emittance compensation

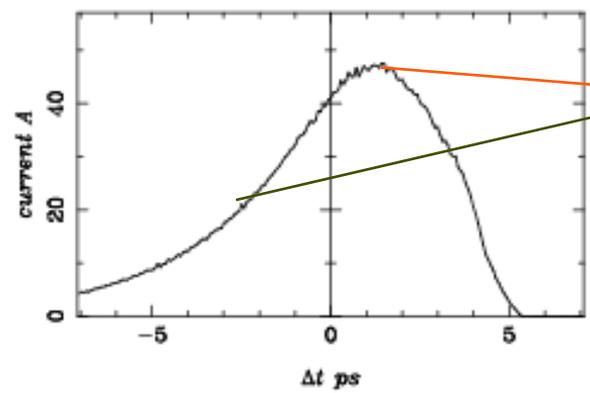
We can compensate for the projected emittance growth by using solenoids



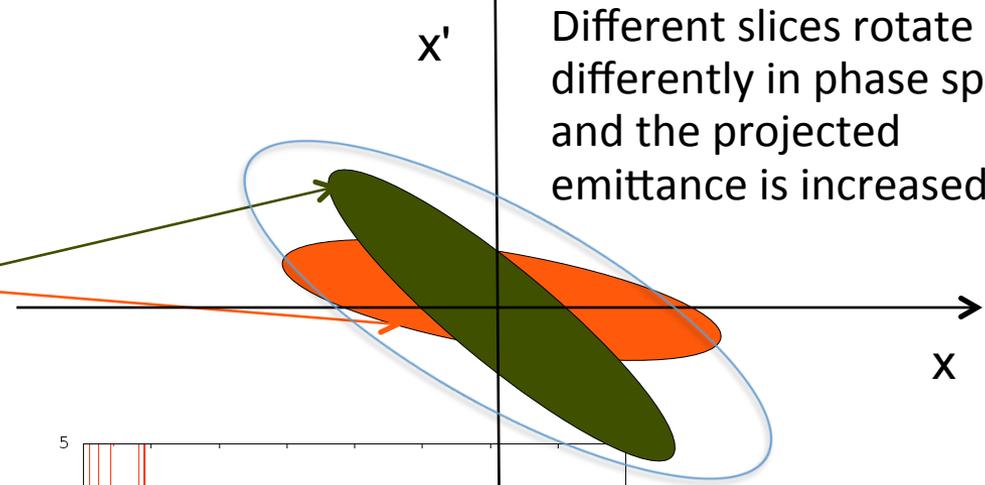
Space charge
Defocusing $K \downarrow j = 2$
 $I \downarrow j / I \uparrow 0 (\beta \gamma) \uparrow 3$

Sol. focusing, $k \downarrow 0 = eB / 2mc\beta\gamma$

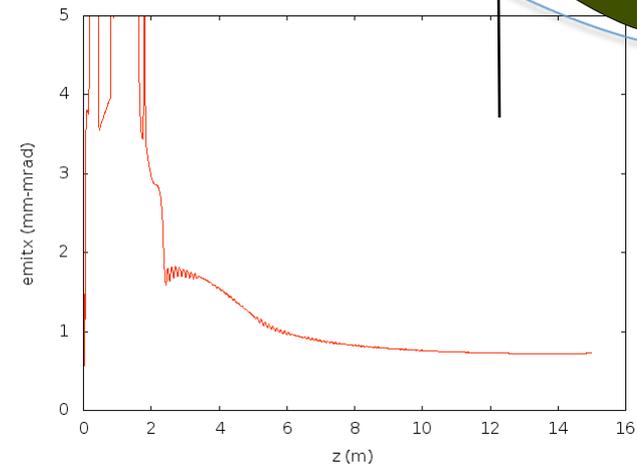
Longitudinal Distribution



Different slices rotate differently in phase space and the projected emittance is increased



To 1st order, the slice ellipses perform small oscillations with different amplitude but the same wavelength. Since they start aligned, they will eventually realign

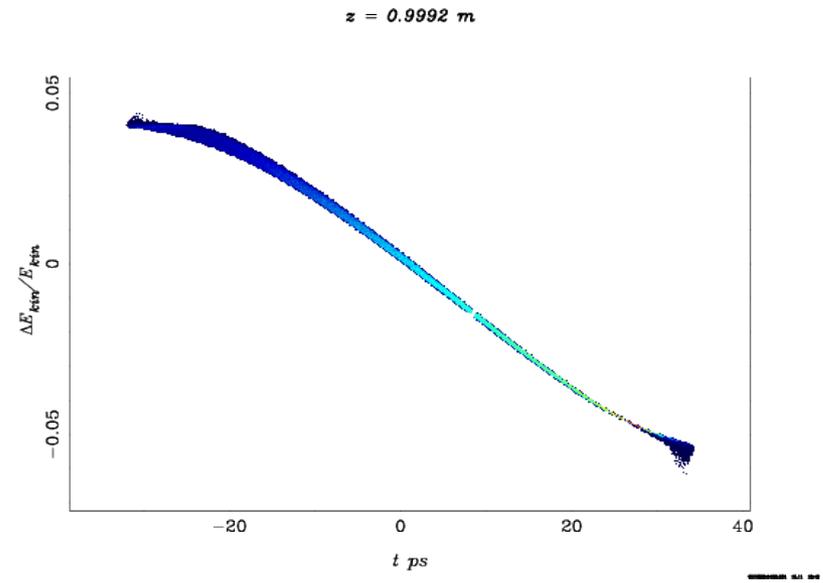
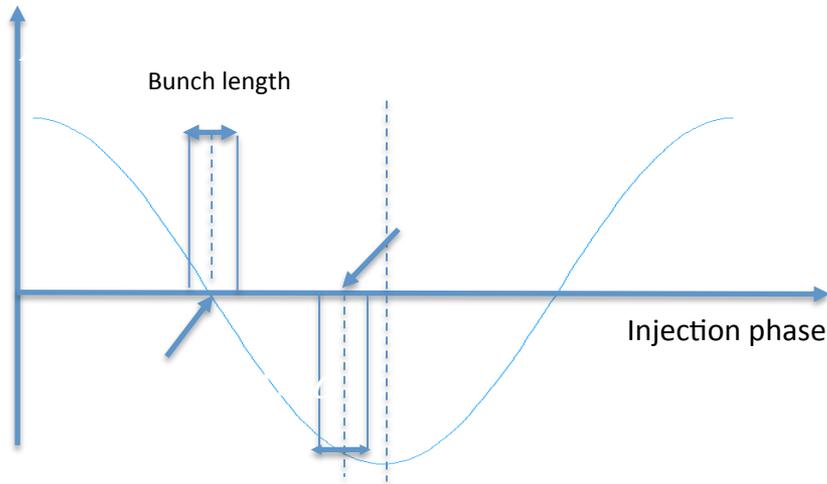


Carlsten 1989
Serafini, Rosenzweig 1997

Compression using RF dephasing

One way to reduce the effect of space charge is to reduce the current at the cathode by increasing the incident laser pulse length.

Simulations show that 5 A produce good beam quality, but we now need to compress in order to get 40-60 A



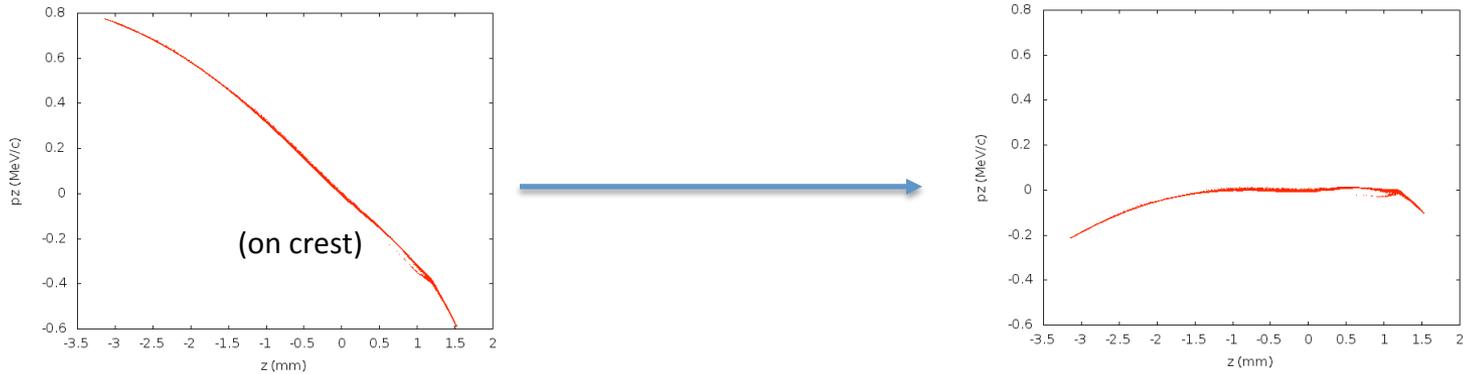
We can compress by injecting at 0 crossing (“ballistic compression”) or with a dephasing , while accelerating (“velocity bunching”). Both are only efficient at low energies since

Ferrario et al 2010

Additional Considerations

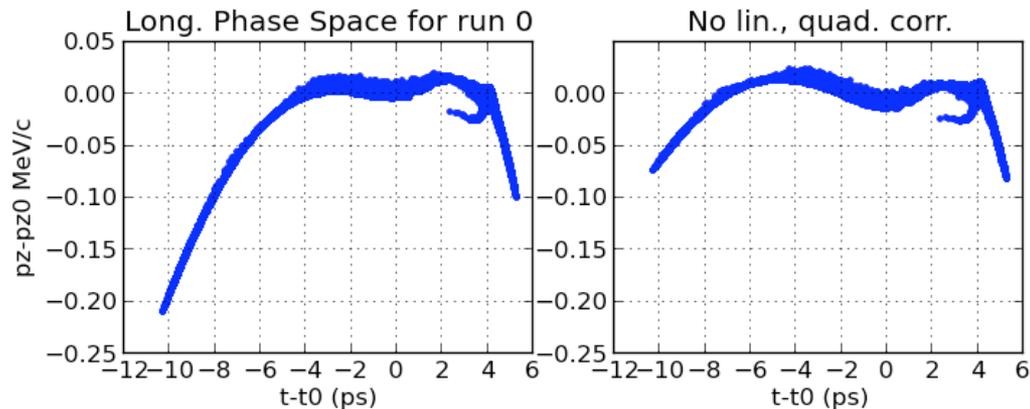
Remove linear correlations (needed for compression) in longitudinal phase space in order to accommodate for the laser heater after the injector.

We can do this by dephasing the last 2 cavities



Simply calculating the rms energy spread overestimates it due to quadratic correlations caused by RF curvature. The 3rd harmonic cavity downstream will remove those.

We mimic that by artificially removing them after the simulation



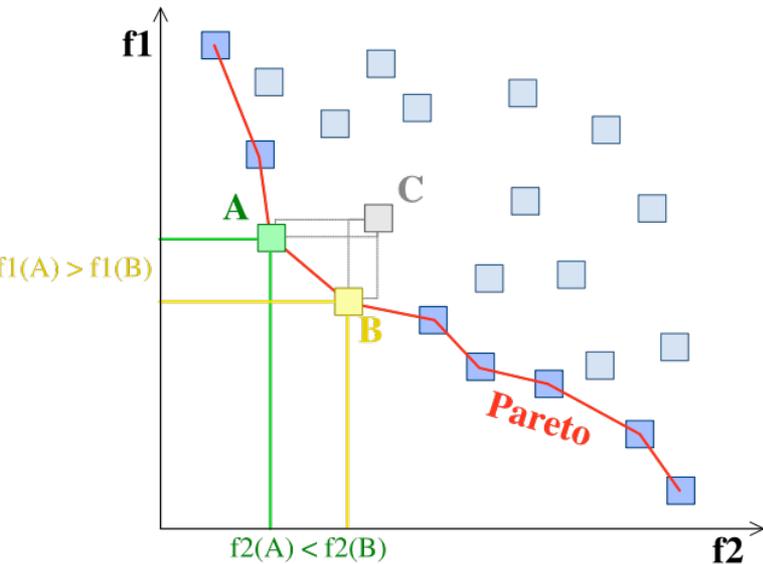
Genetic Optimizer

The problem:

Find global optimum(s) for a problem with multiple, non-linearly coupled knobs (typically more than 10)

The solution:

Multi-Objective Genetic Algorithms



Vilfredo Pareto
1848-1923

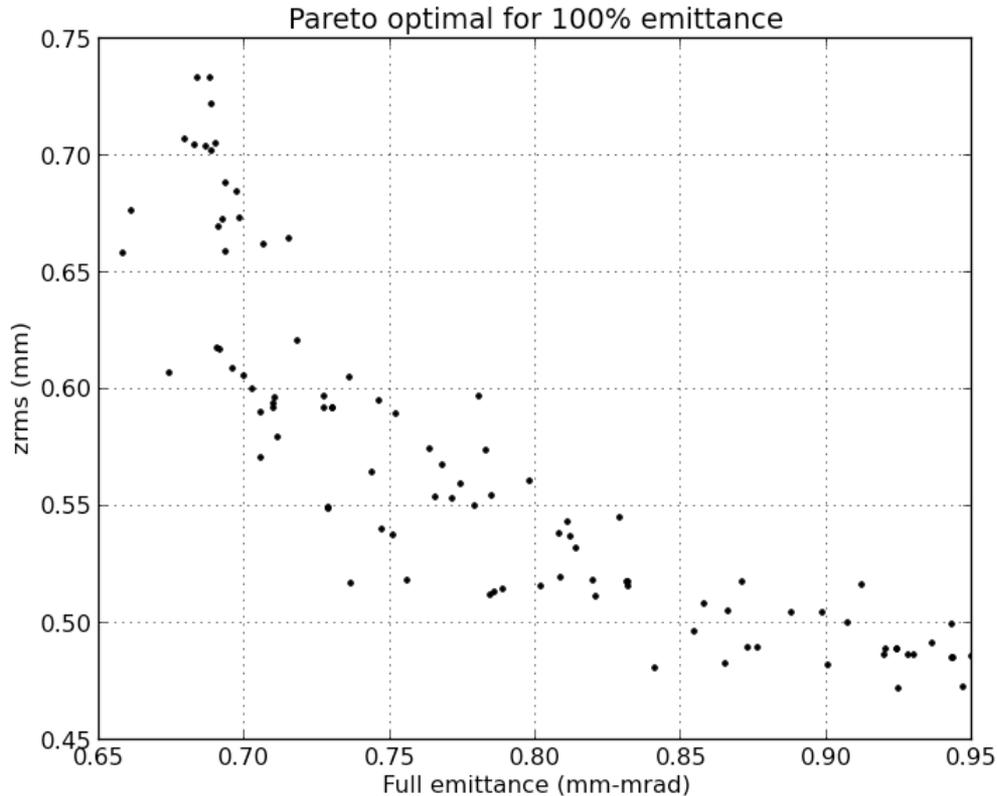
We use the algorithm NSGA2
(Deb 2002, Bazarov 2005)

Parallelized evaluation of the solutions
(the most computationally expensive part).
Runs performed in the alsacc cluster.

A and B dominate C
But A and B don't dominate each other
The algorithm finds a front of non-dominated solutions

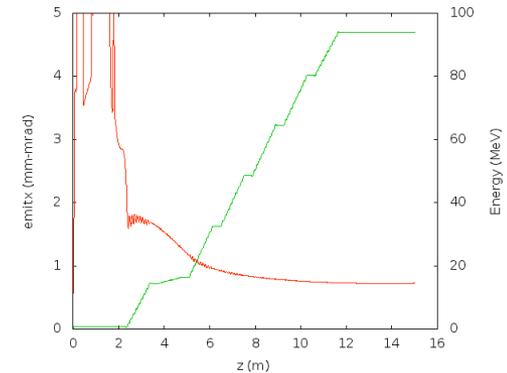
Photos: wikipedia

Pareto Optimum



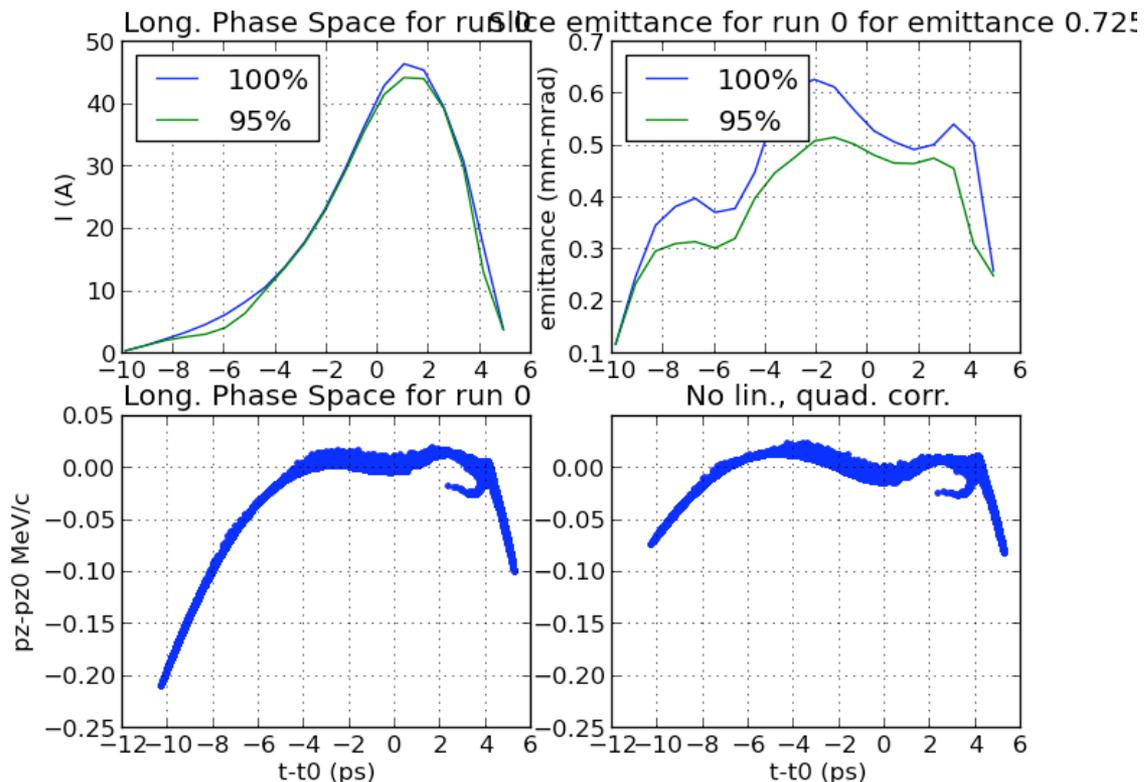
Simulations performed using ASTRA

Emittance vs bunch length at the injector exit (~100 MeV)
Space charge effects should be “frozen-in” at this point



The result is a front of solutions
Using this, we can evaluate the trade-offs and choose 1 (or possibly more) as candidates

Optimized solution

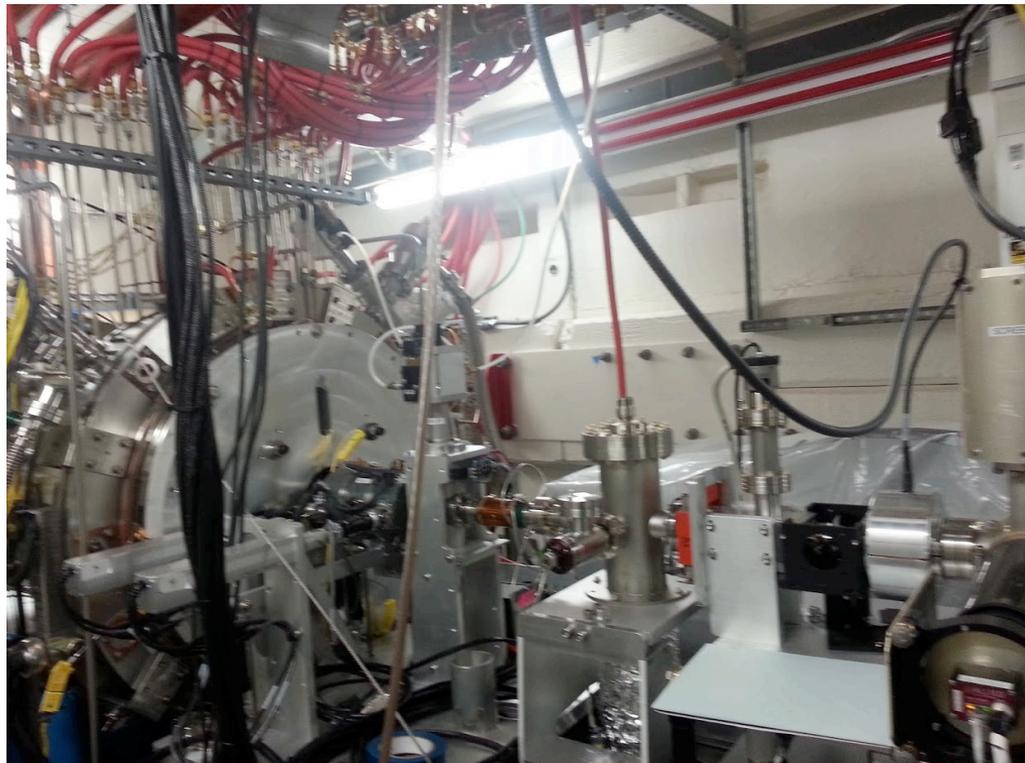


Gun Phi (deg)	-2.91674
Buncher Grad. (MV/m)	3.99863
Buncher Phi (deg)	-71.1609
Sol. 1 B (T)	0.0396057
Sol. 2 B (T)	0.0282948
Cav. 1 Grad. (MV/m)	29.4731
Cav. 1 Phi (deg)	-40.6049
Cav. 2 Grad. (MV/m)	6.18249
Cav. 2 Phi (deg)	-52.2731
Sig x (mm)	0.304257
Laser Δt (ns)	0.0440136
Cav. 6 Phi (deg)	13.1272
Cav. 7 Phi (deg)	33.4975

Final beam Characteristics:
 100% emitx = 0.725 mm-mrad
 95% emitx = 0.667 mm-mrad
 $I \sim 45$ A
 energy = 93.88 MeV
 cor. e-spread = 20.1 keV
 sigx final = 196 μm

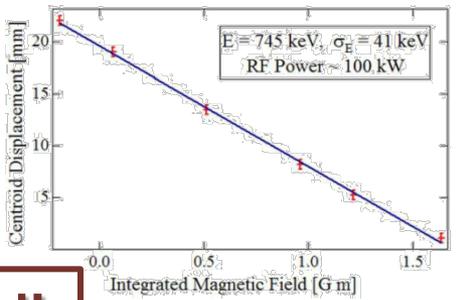
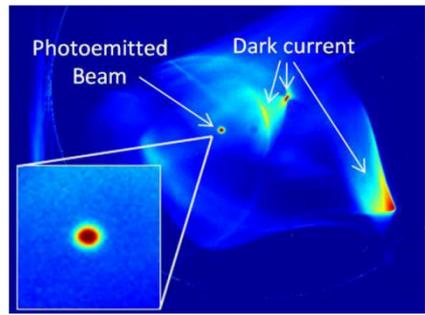
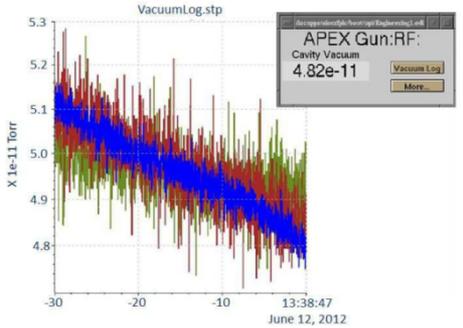
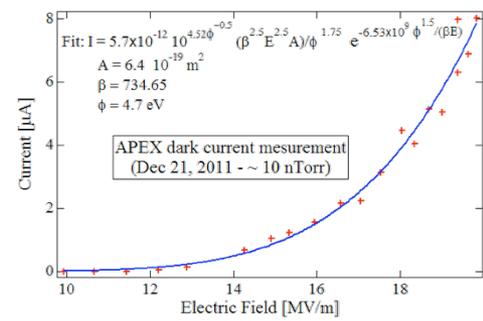
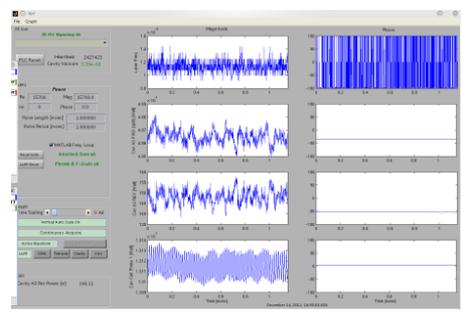
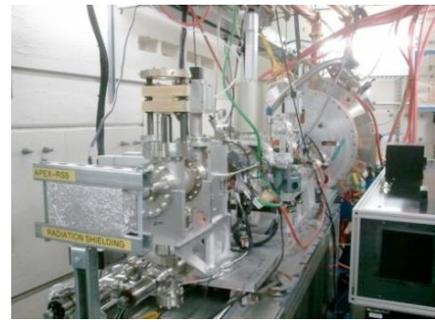
Start-to-end simulations are currently under way

Advanced Photoinjector EXperiment



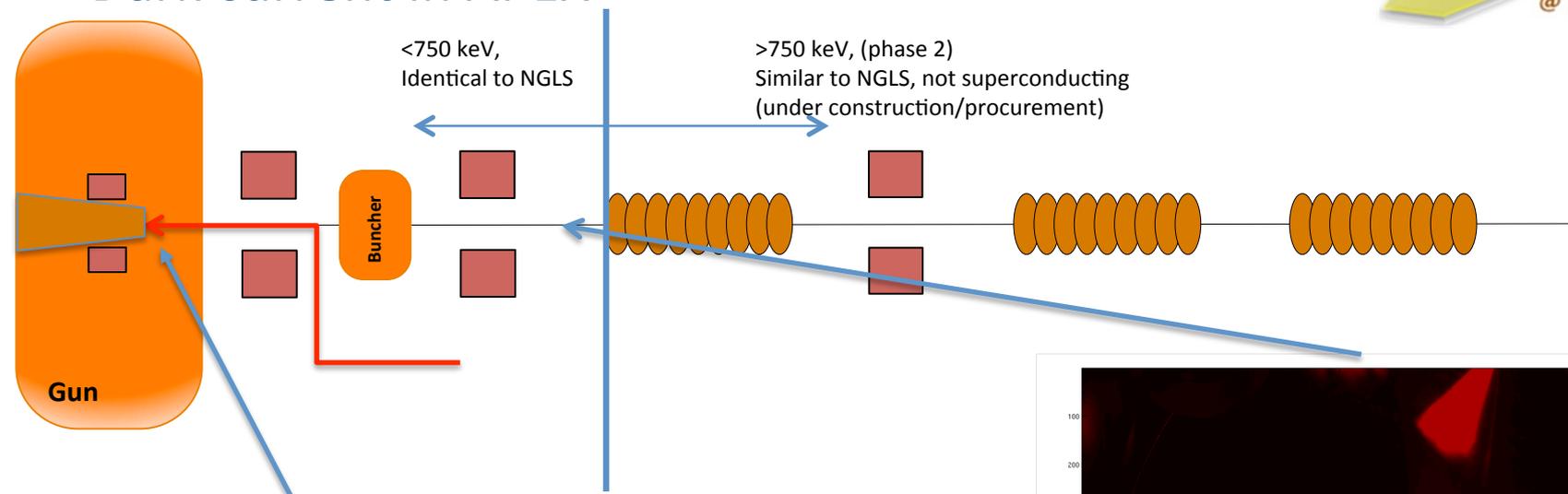
APEX PRESENT ACHIEVEMENTS

- Gun and Phase 0 beamline built and installed.
- Gun conditioned at full RF power.
- Dark current characterized: $< \sim 1 \mu\text{A}$ (acceptable level)
- Very promising vacuum performance from initial tests.
- First photo-emitted electron beam with low QE cathode generated.
- Nominal photo-emitted electron beam energy demonstrated.

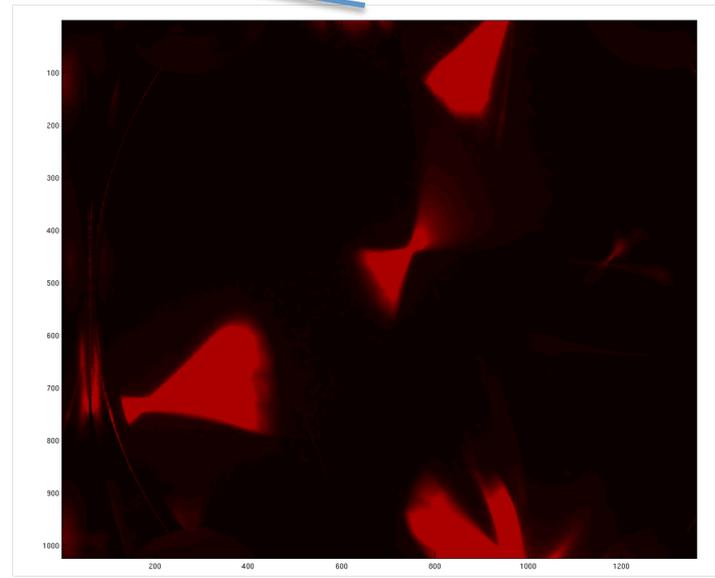
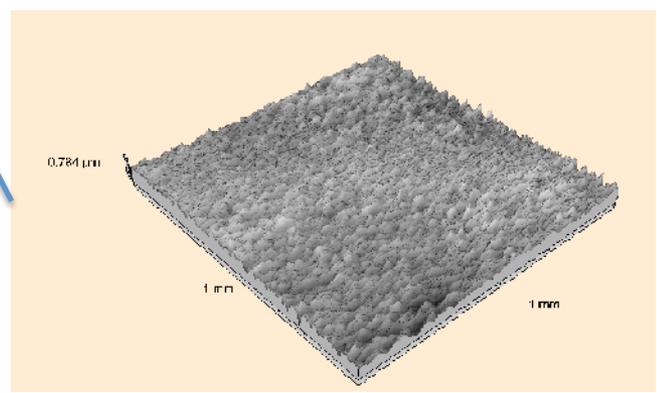


Gun technology fully demonstrated!
F. Sannibale, D. Filippetto, C. F. Papadopoulos
et al., PRST-AB 15, 103501 (2012)

Dark current in APEX



- Solenoid magnet
- Single-cell RF cavity
- Multi-cell RF cavity
- Laser pulse



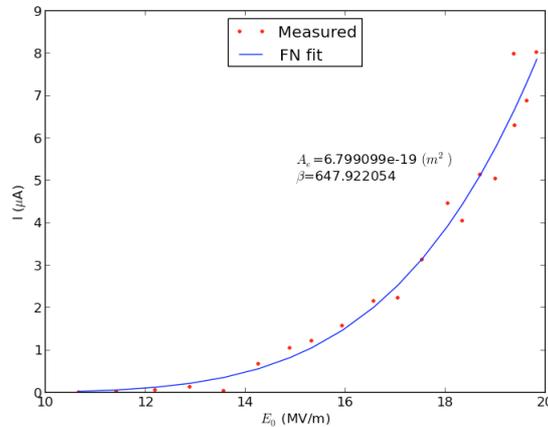
Cathode surface, similar to APEX/NGLS
 Paolo Michelato, INFN Milano – LASA
 “High QE Photocathodes lifetime and dark current investigation”,
 PITZ collaboration meeting 4-5.02.006

Dark current “Hotspots”
 Butterfly shape due to large energy spread
 D. Filippetto, F. Sannibale

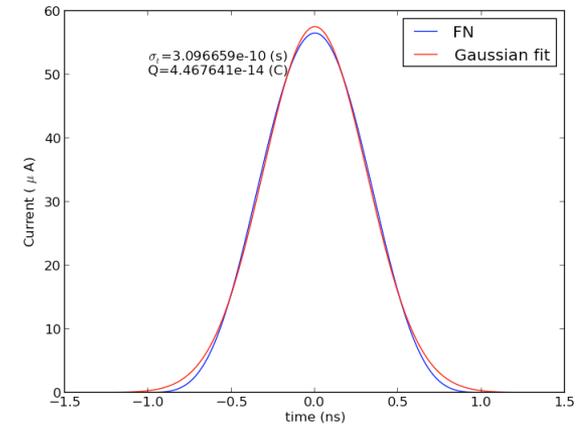
Fowler-Nordheim model of Dark current emission

Use integrated Fowler - Nordheim fit to calculate pulse with instantaneous $E(t) = E_0 \cos(\omega t)$

Dark Current measurements
F. Sannibale



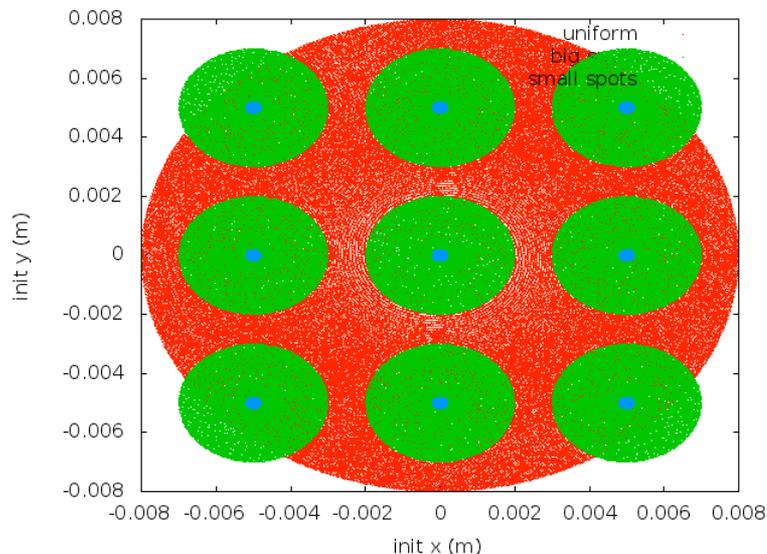
Dark Current profile



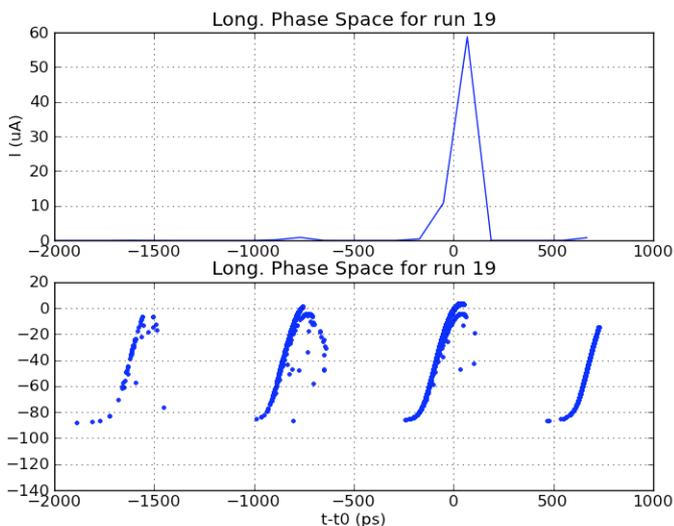
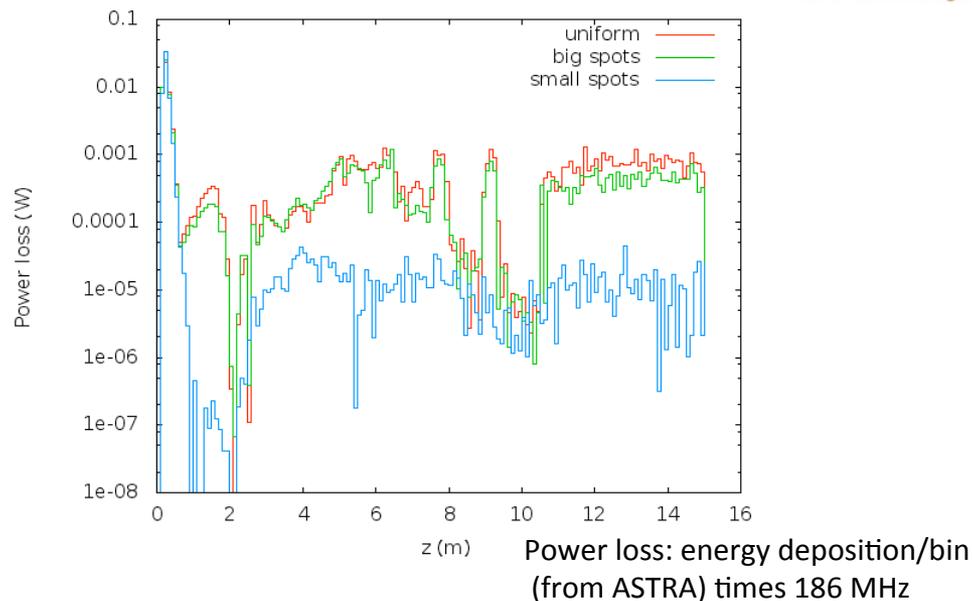
J. W. Wang, G. A. Loew
SLAC PUB 7684

Dark Current:
44 fC*186 MHz \approx 8.1 μA
Nominal Beam (average):
300 pC*1MHz \approx 300 μA

Dark current transport



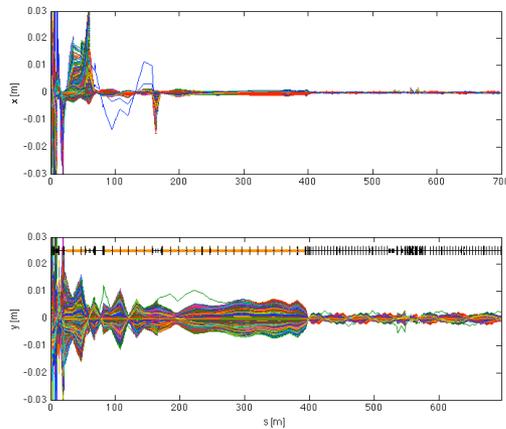
Simulate multiple transverse distributions



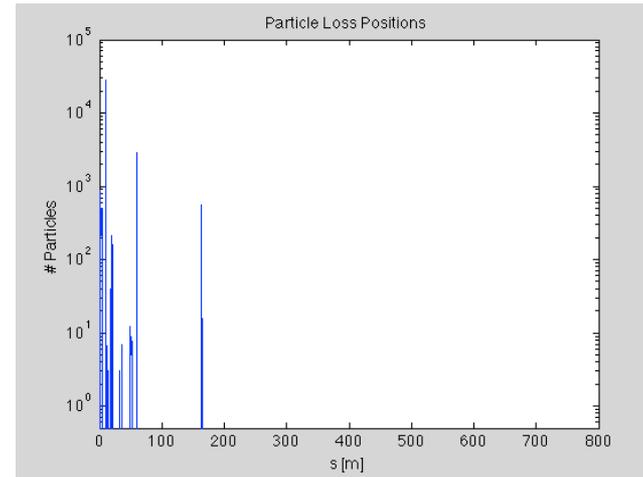
In all cases
~ 85% charge is lost in injector

Longitudinal Phase Space at
Injector exit: The dark current
pulse is split due to different
RF wavelengths

Dark current collimation schemes

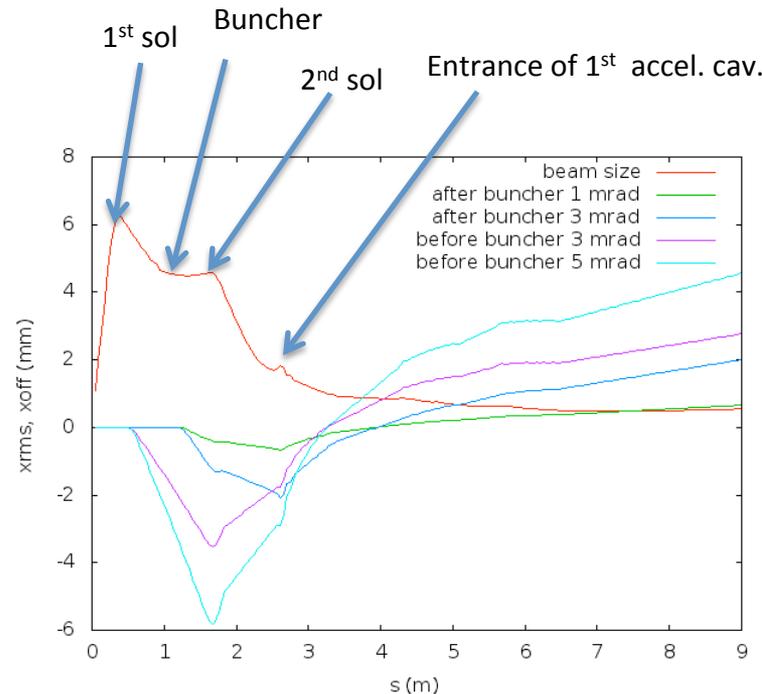


Linac collimators
(C. Steier)



Initial collimation simulations for the injector.

Use a modified version of the ALS kicker cavity to kick the dark current
C.F. Papadopoulos, H. Qian

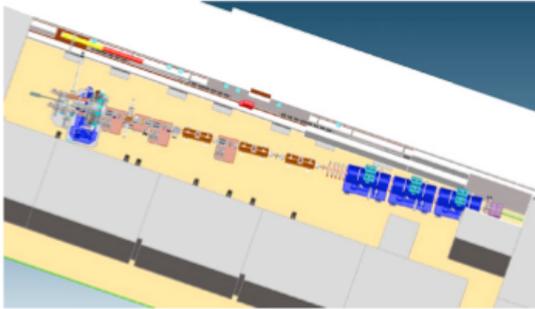
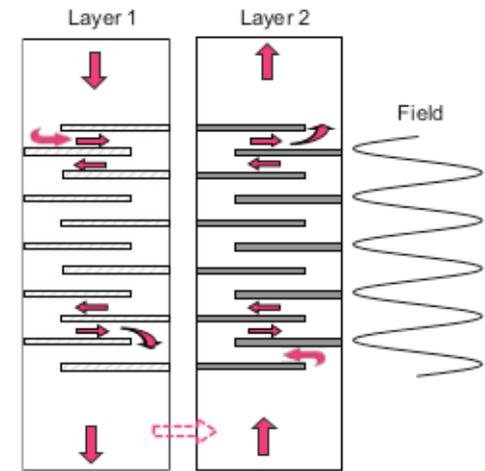


BTFEL: A proposed seeded IR FEL amplifier

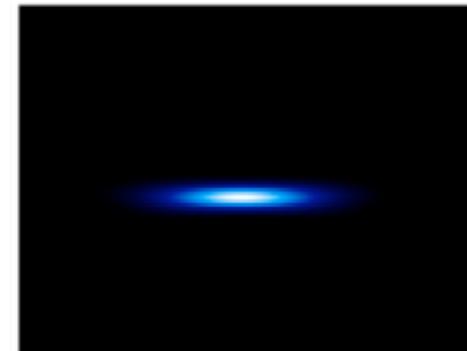
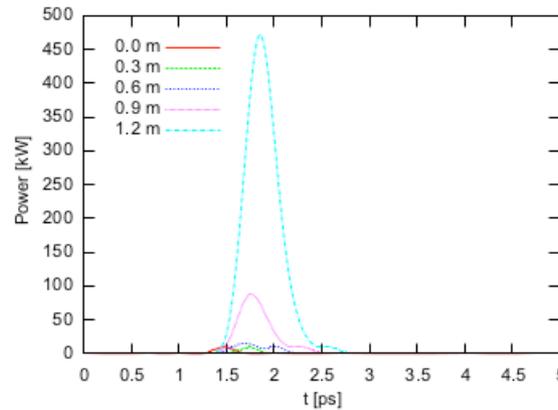
Use the APEX beam
to drive an FEL.
(50 pC, 23 MeV, 40 A)

Use (and test)
a superconducting
tape undulator

Undulator param.



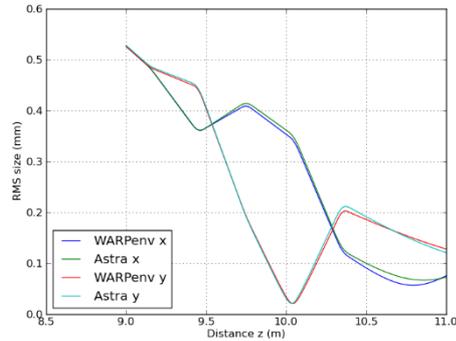
In order to produce a
seeded,
single spike pulse



Space charge effects in the BTFEL

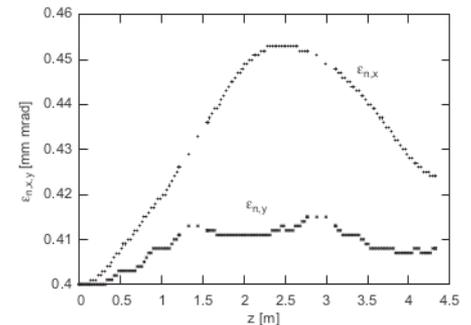
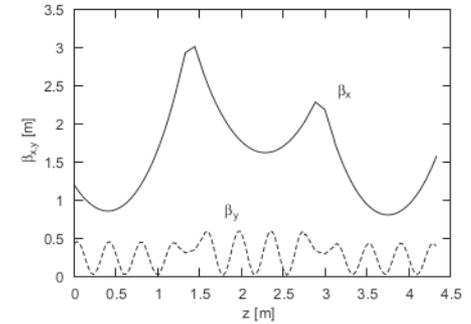
Due to the low energy and the relatively high current, space charge effects are important in the BTFEL

Take care to match with SC in the beamline



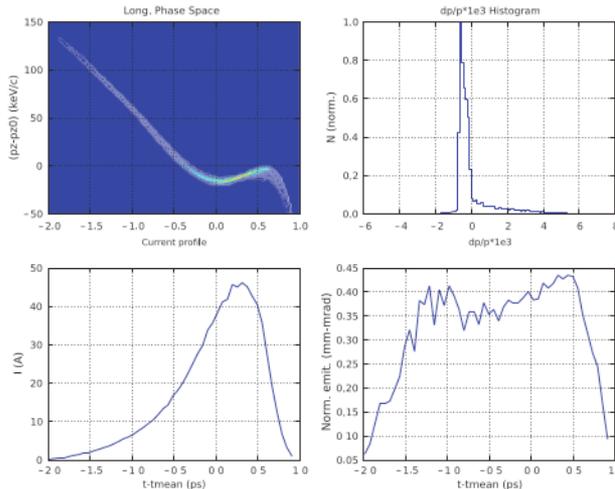
Quadrupole matching in the beamline

As well as in the undulator



Emittance oscillation in the undulator

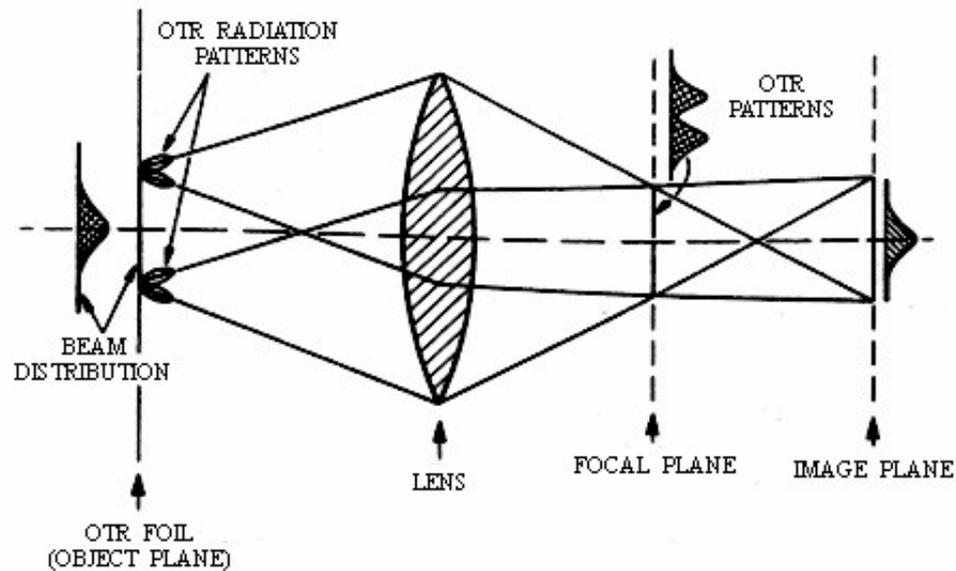
But space charge can also help to remove (some) of the chirp without dephasing and losing energy



Yoon, M.; Filippetto, D.; Papadopoulos, C. F.; Pellegrini, C.; Penn, G.; Prestemon, S.; Sannibale, F.

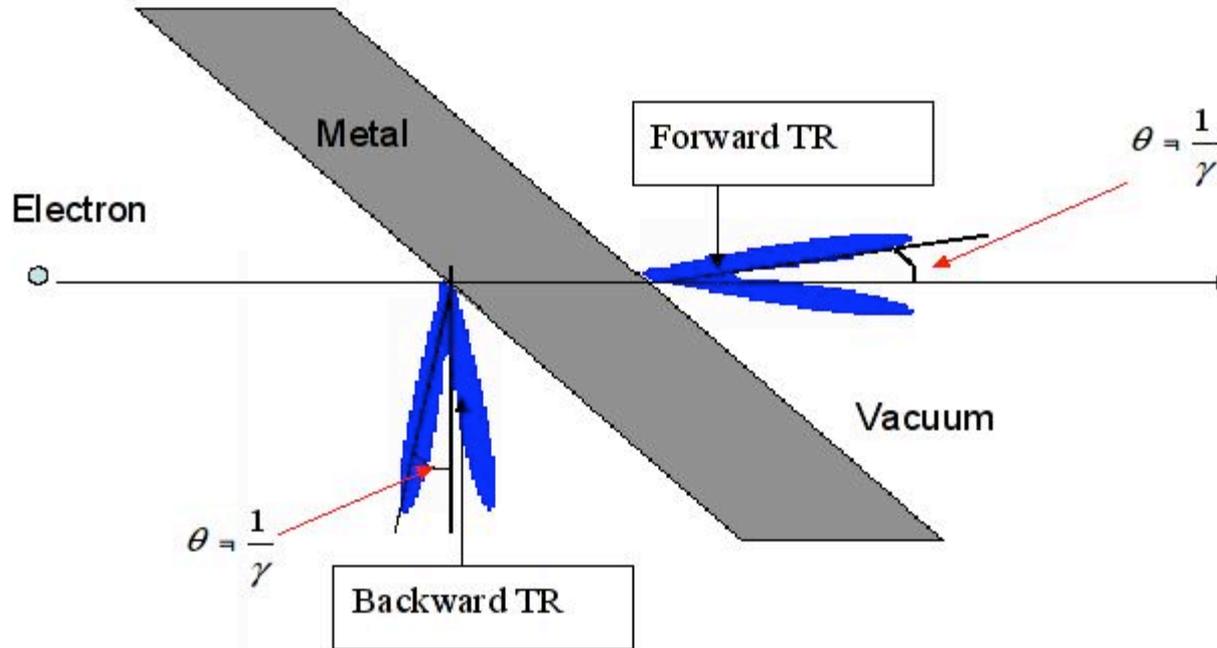
NIMA, Volume 660, Issue 1, p. 138-146.

Measuring emittance using transition and diffraction radiation



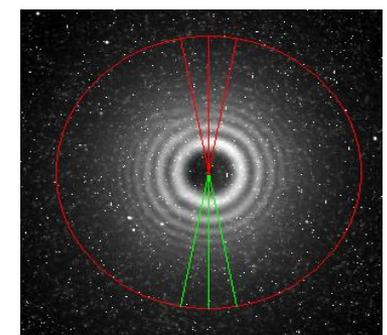
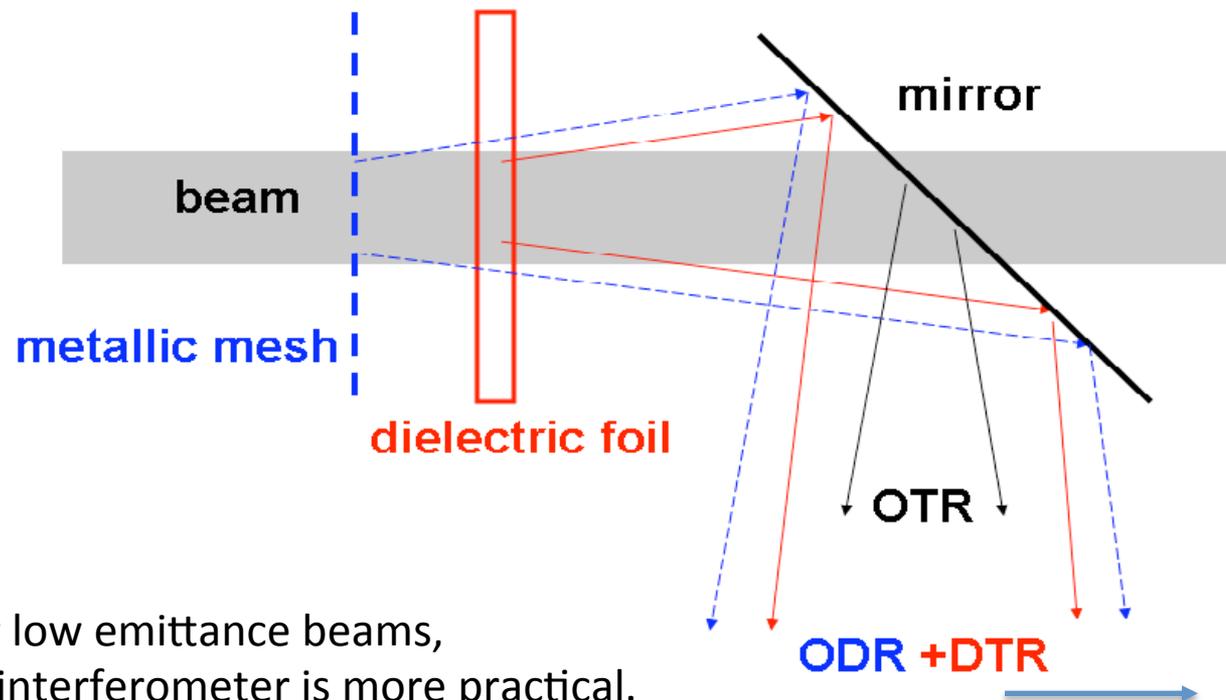
Source: R. Fiorito

Angular dependence of Optical Transition Radiation



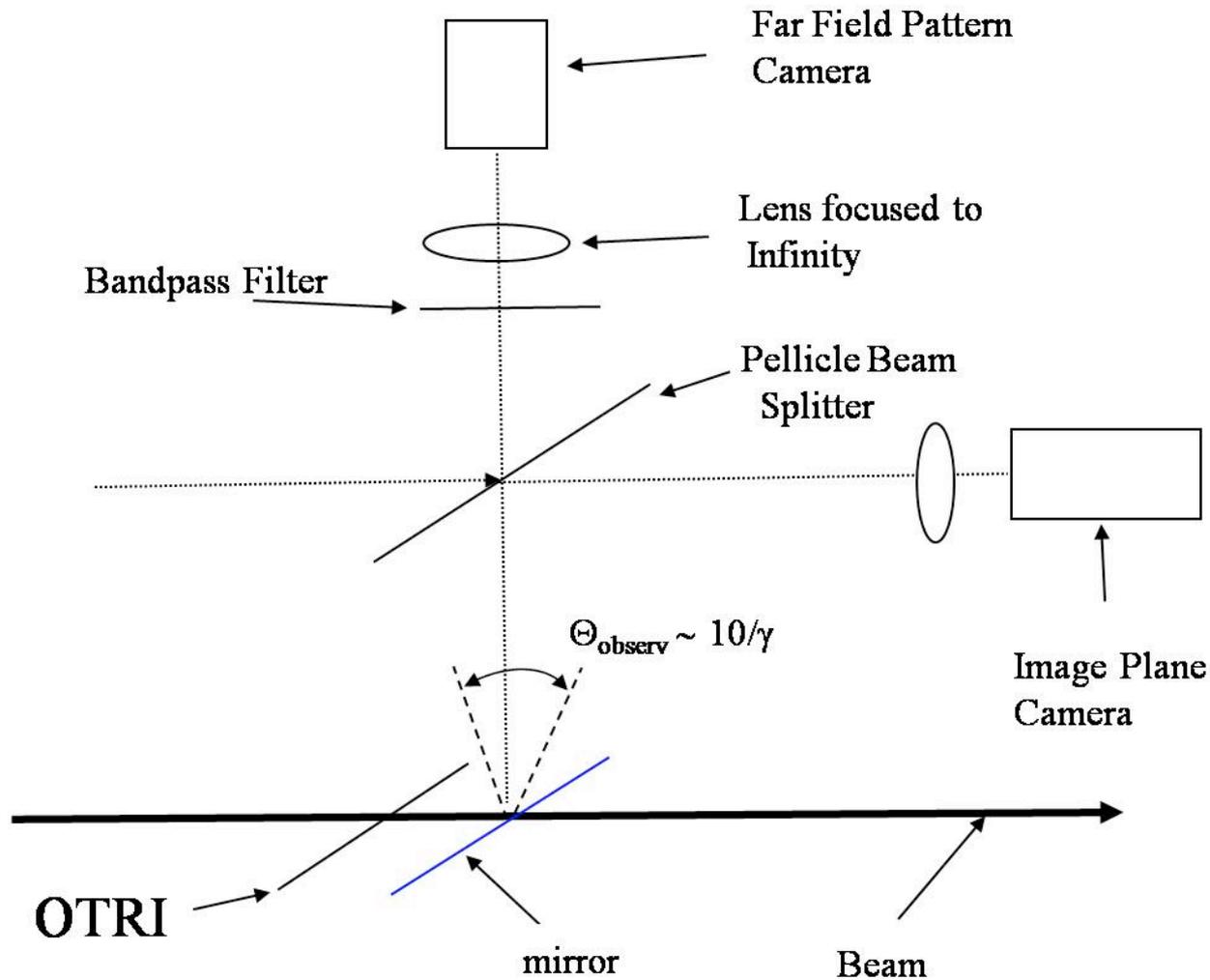
The single particle OTR angular distribution contains information about the incidence angle of the incoming electron. The same holds for diffraction radiation etc

Optical diffraction radiation - Dielectric transition radiation interferometry



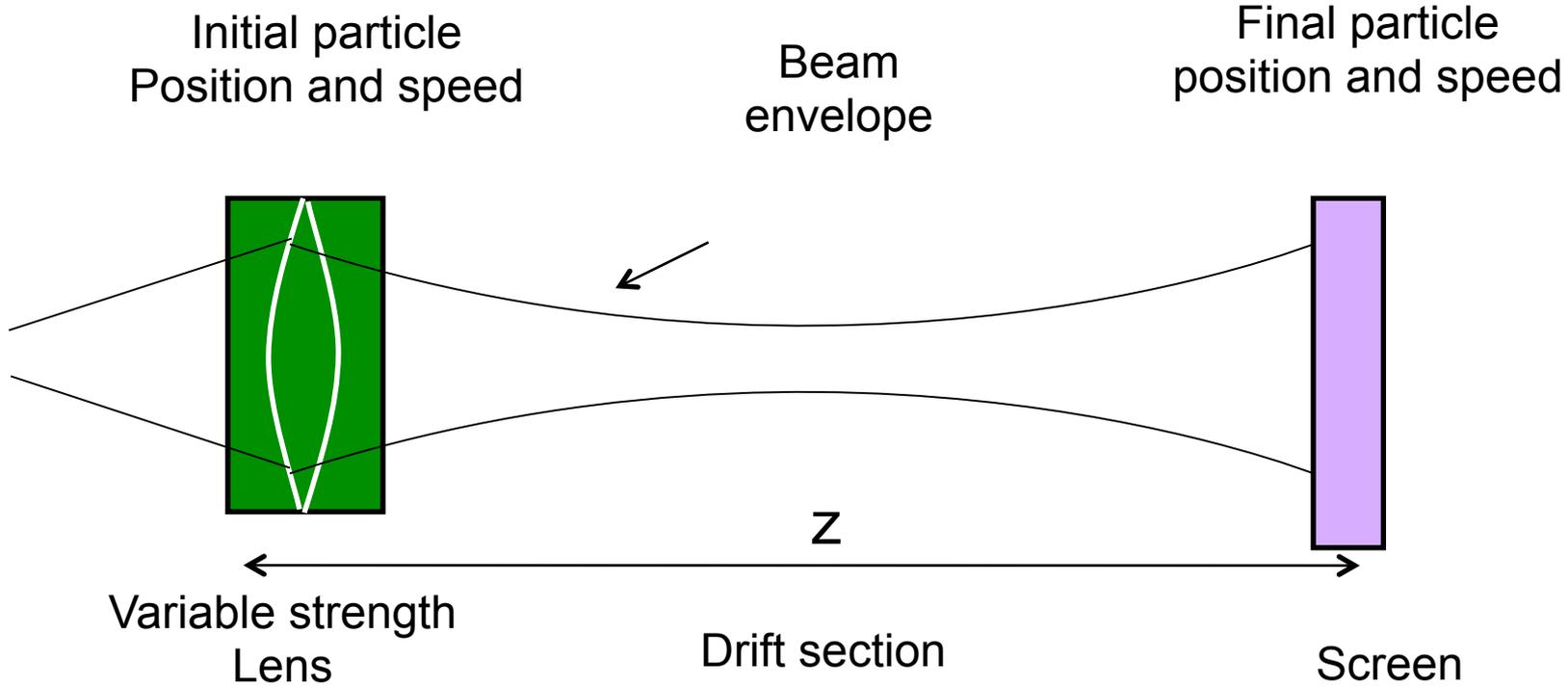
For low emittance beams, an interferometer is more practical. We can convolute the single electron interference pattern with the angular distribution of all the particles, for example a Gaussian. The divergence of the beam can then be calculated from the visibility of the fringes, if we know the energy spread.

Experimental setup

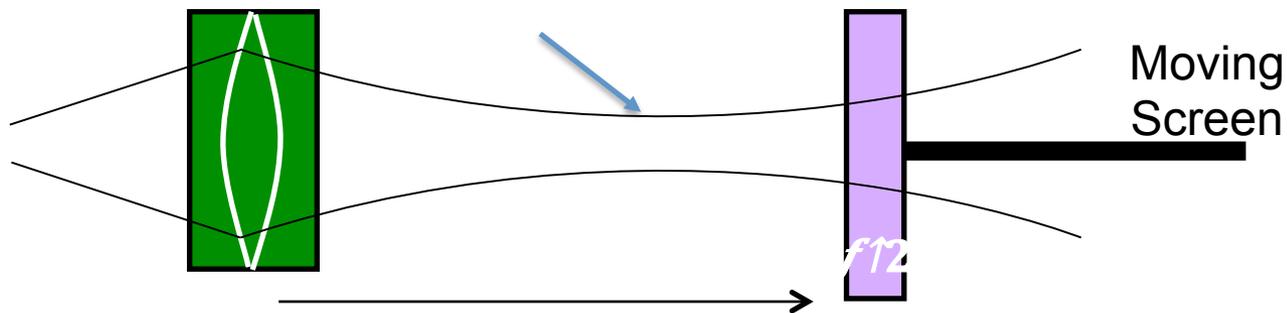


OTR Interferometry allows simultaneous measurement of beam size and beam divergence

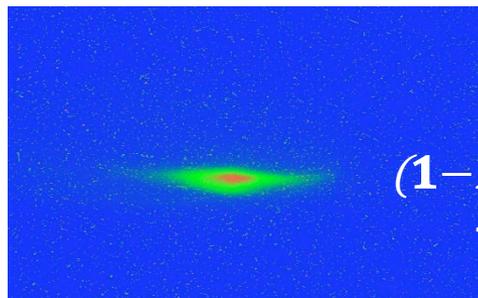
Thin lens scan



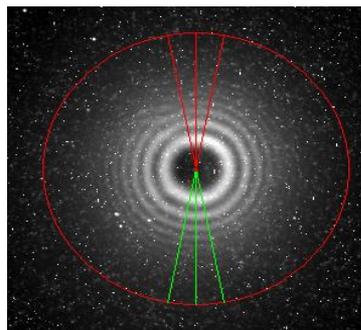
Different experimental conditions for OTR measurements



Keep the screen at a constant position, vary the magnet



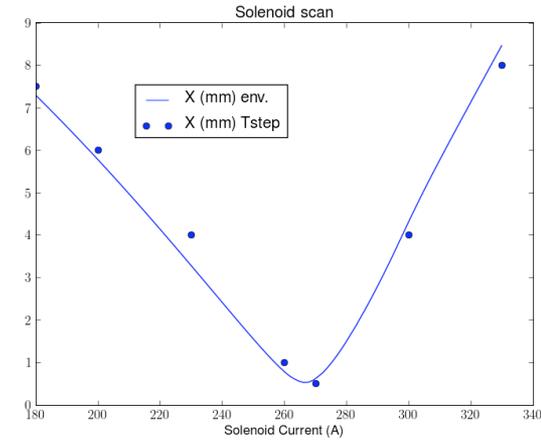
Minimize the size in x (or y)



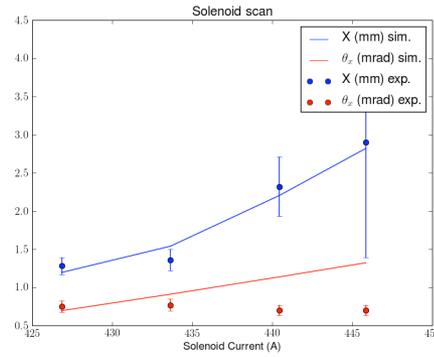
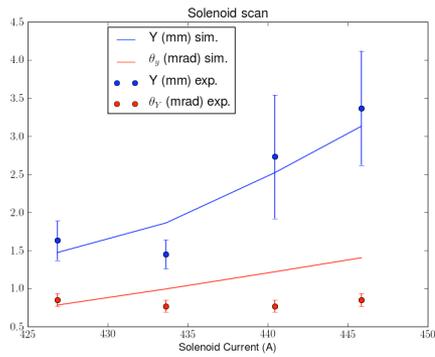
Maximize the visibility of the fringes

Measurements at AWA

- Develop an envelope code (with SC and RF focusing effects included) to model the soleoid scan
- Using the fast envelope code, we scan the 3 dimensional parameter space ().
- The emittance value that minimizes χ^2 is 19 mm-mrad for x and 22 mm-mrad for y



Comparison of envelope code with Tstep



Extensions of the method may include different radiation types or energy spread measurements

Solenoid Current (A)	X size (mm)	Θ_x (mrad)	Y size (mm)	Θ_y (mrad)
427	1.28	0.75	1.63	0.85
434	1.36	0.77	1.45	0.77
440	2.32	0.7	2.73	0.77
445	2.9	0.7	3.37	0.85

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